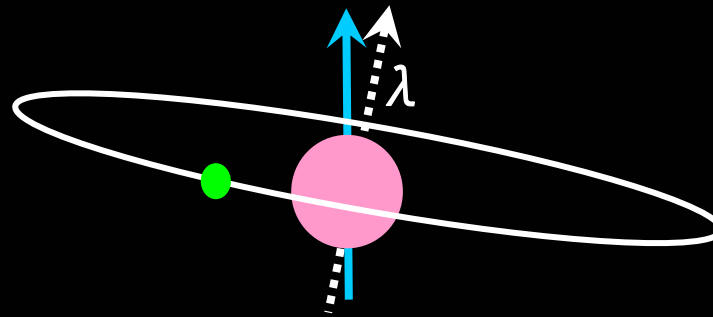
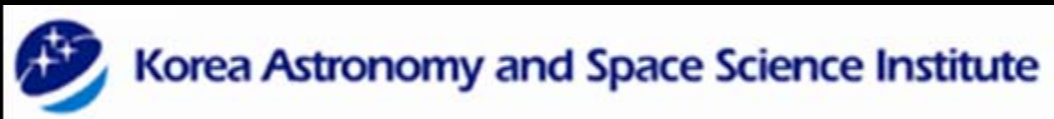


Transiting extra-solar planets: spin-orbit misalignment and rings



Yasushi Suto *Department of Physics, University of Tokyo*



KASI colloquium
September 26, 2006

Recent Activities of Observational Cosmology Group, University of Tokyo (1)

■ SDSS galaxy and quasar statistics

- genus statistics and phase correlation of SDSS galaxies (Hikage et al. 2003, 2004, 2005; Hikage, Matsubara, and Suto 2004; Park et al. 2005)
- 3pt correlation functions of SDSS galaxies (Kayo, Suto, Nichol et al. 2004; Nichol et al. 2006)
- widest-separation lensed quasar from SDSS (Inada et al. 2003; Oguri et al. 2004; Ota et al. 2006)
- 2pt correlation functions of SDSS quasars and cosmological constant (Yahata et al. 2005)
- constraints on the deviation from Newton's law of gravity from SDSS galaxy power spectrum (Shirata, Shiromizu, Yoshida & Suto 2005)
- SDSS galaxy number counts and SFD dust map (Yahata et al. [astroph/0607098](#))

Recent Activities of Observational Cosmology Group, University of Tokyo (2)

■ Dark halo and galaxy cluster

- highest-angular resolution SZ maps in submm and mm (Komatsu et al. 1999, 2001; Kitayama et al. 2004)
- triaxial modeling of dark matter halos (Jing & Suto 2002; Oguri, Lee & Suto 2003; Lee, Jing & Suto 2005)

■ Warm/hot intergalactic medium (WHIM)

- a proposal of oxygen emission line search with DIOS (Yoshikawa et al. 2003, 2004)
- feasibility of an absorption line search with XEUS along bright quasars and GRB afterglow (Kawahara et al. 2006)

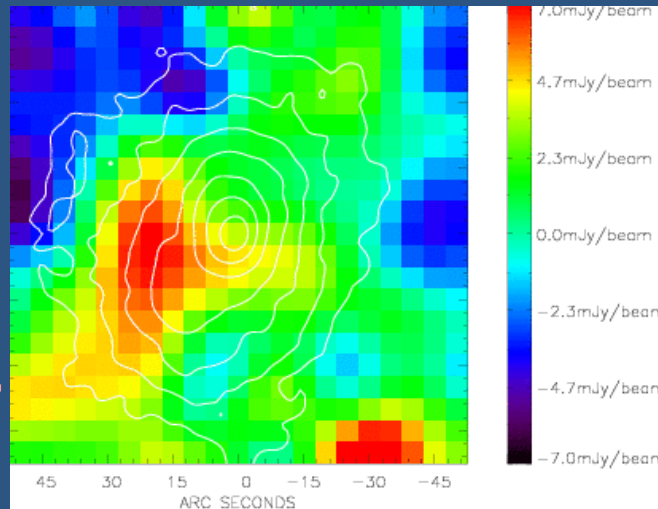
■ Spectroscopy of transiting extrasolar planets

- constraints on planetary atmosphere (Winn et al. 2004; Narita et al. 2005)
- first detection of the spin-orbit misalignment in an extrasolar planetary system using the Rossiter effect (Ohta, Taruya & Suto 2005; Winn et al. 2005)

Observation of the Sunyaev-Zel'dovich effect

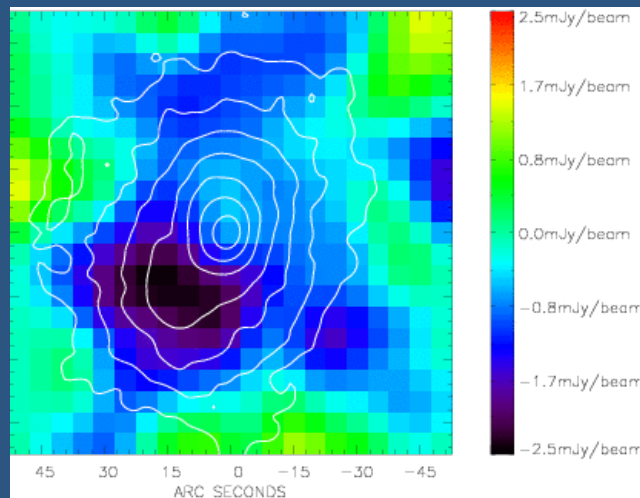
the most luminous X-ray cluster: RX J1347.5-1145

The *first*
submm SZ
map with
SCUBA, JCMT

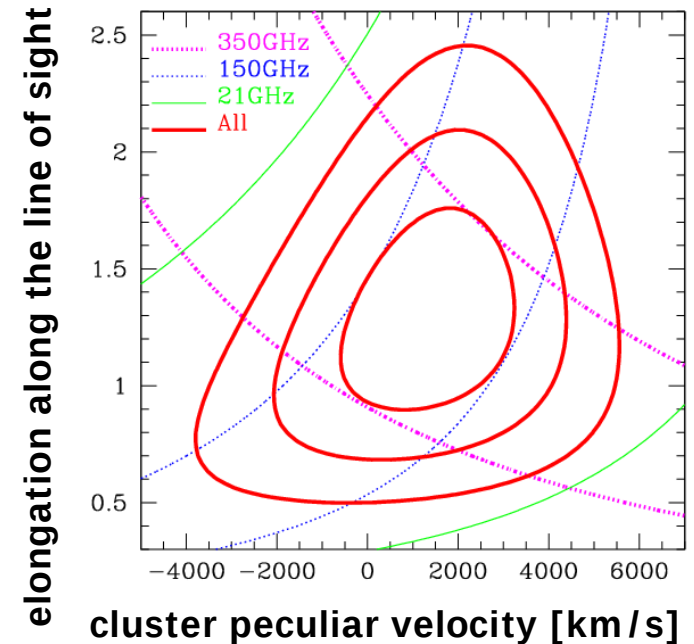


submm SZ map

The *highest*
angular
resolution
($\sigma_{\text{FWHM}} = 13''$)
mm SZ map
with NOBA,
Nobeyama



mm SZ map: discovery of substructure



Komatsu et al.

ApJ 516 (1998) L1

PASJ 53 (2001) 57

Kitayama et al.

PASJ 56 (2004) 17

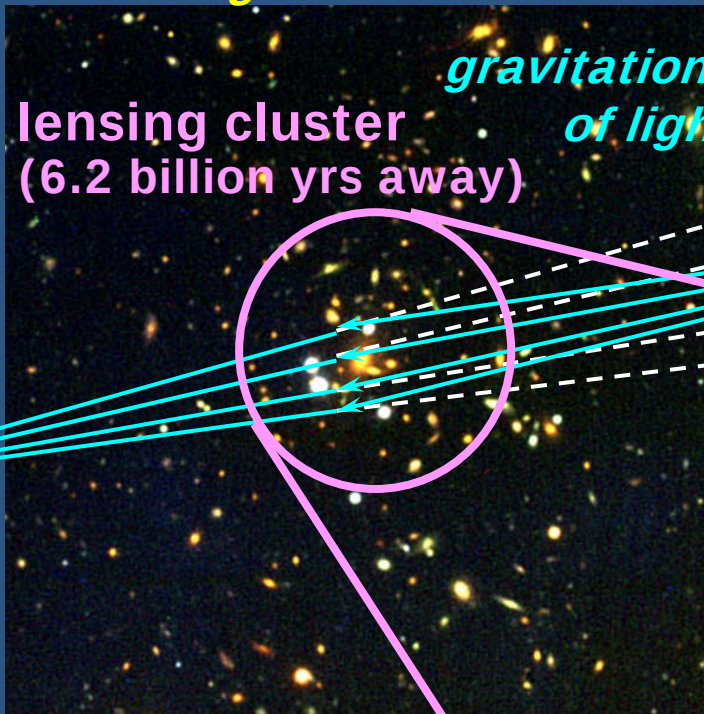
Widest-separation lensed quasars

Inada, Oguri et al. (2003)

Subaru 8.2m telescope



Subaru image



lensing cluster
(6.2 billion yrs away)

gravitational bending
of light path

image 1

image 2

SDSS J1004
(9.8 billion yrs away)

image 3

image 4

Zoomed Subaru image



1

2

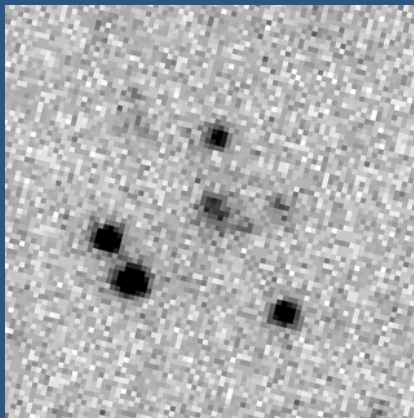
3

4

410,000 light yrs
apart



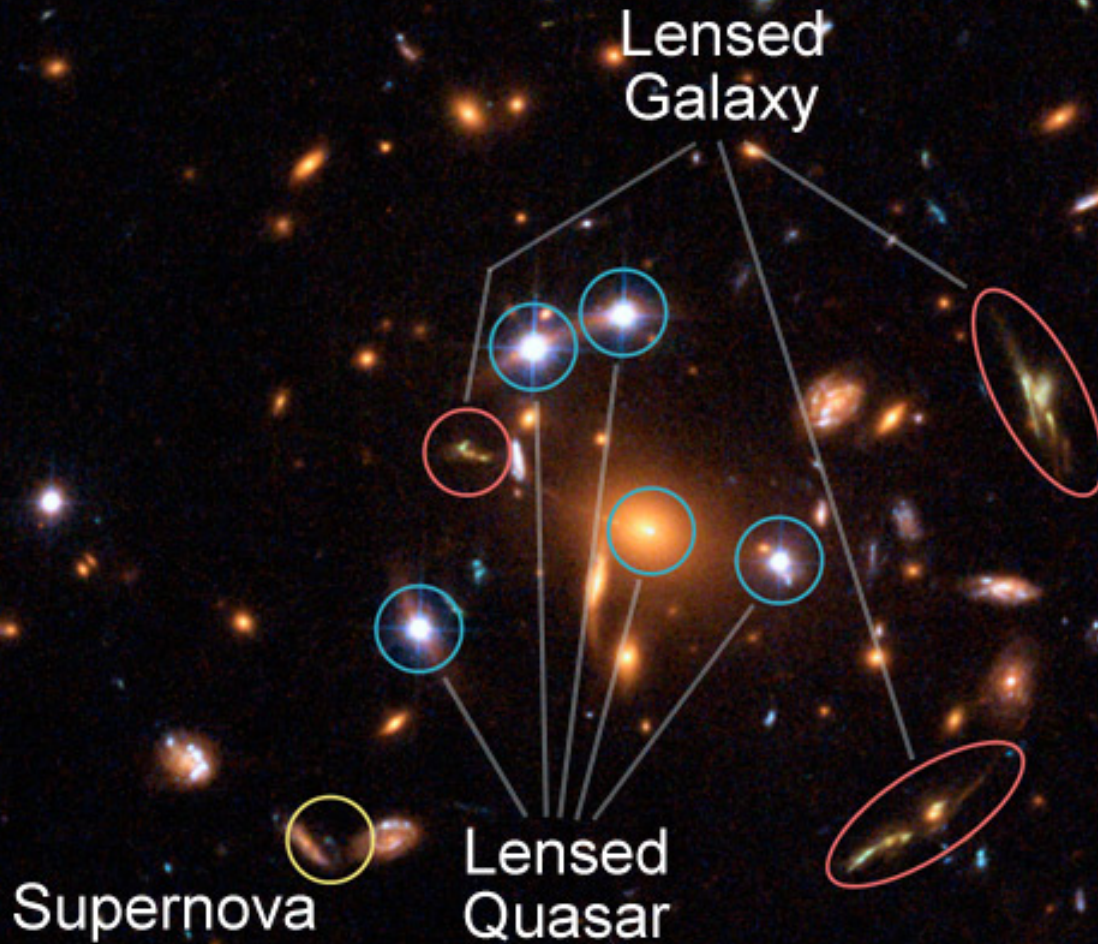
SDSS 2.5m telescope



SDSS image

Galaxy Cluster SDSS J1004+4112
HST ACS/WFC

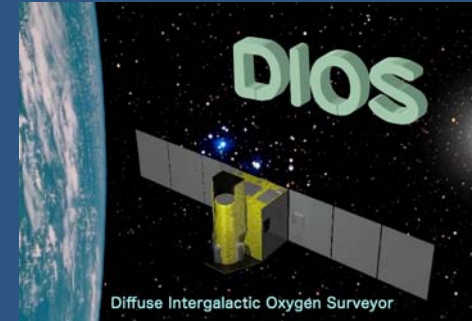
SDSS J1004+4112



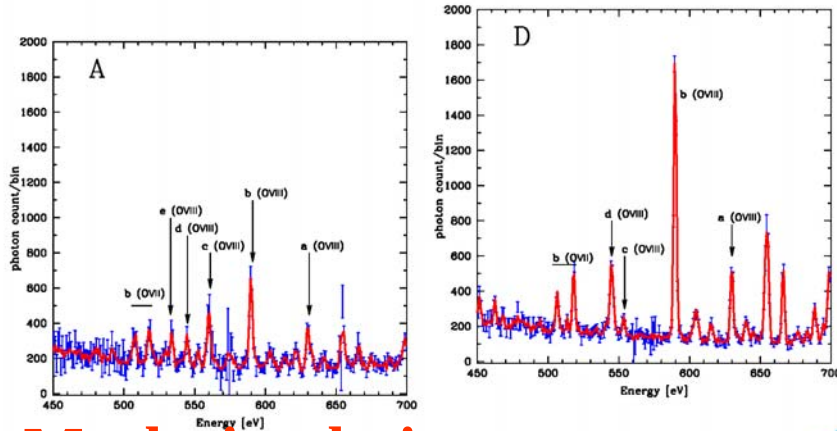
HST photo-release on May 23, 2006

<http://hubblesite.org/newscenter/newsdesk/archive/releases/2006/23/>

Searching for cosmic missing baryons with DIOS (Diffuse Intergalactic Oxygen Surveyor)



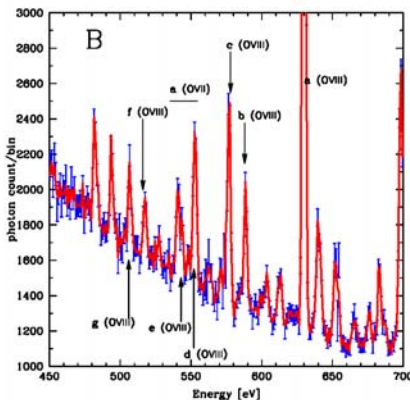
Yoshikawa et al.
PASJ 55 (2003) 879
PASJ 56 (2004) 939



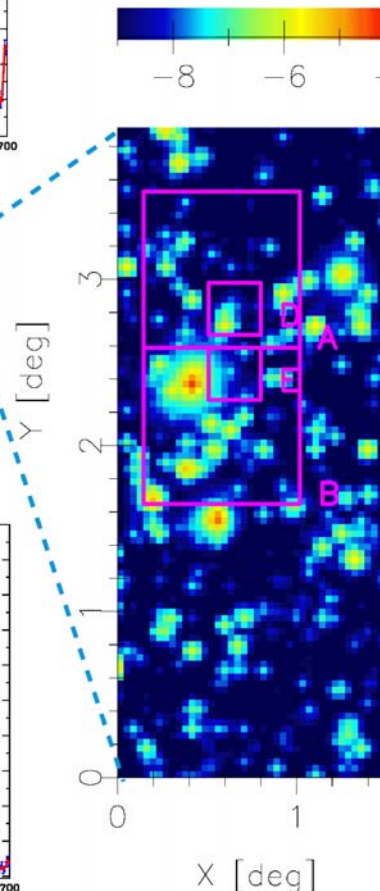
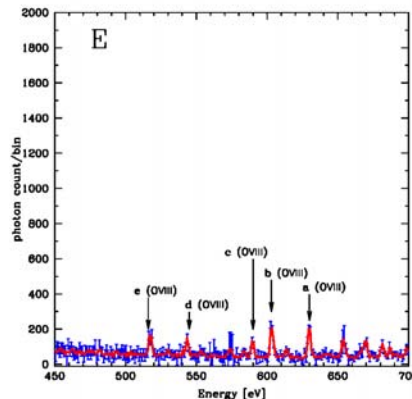
Mock simulations



$z=0$



$z=0.3$



Tokyo Metropolitan Univ.:

T. Ohashi

JAXA/ISAS:

N. Yamasaki

K. Mitsuda

Nagoya Univ.:

Y. Tawara

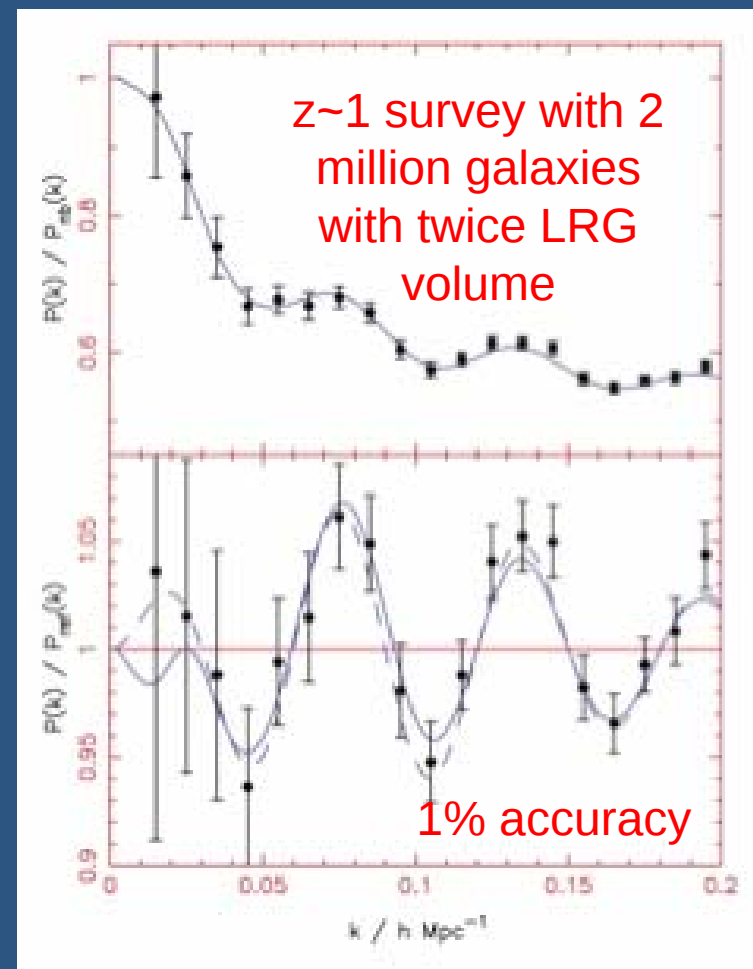
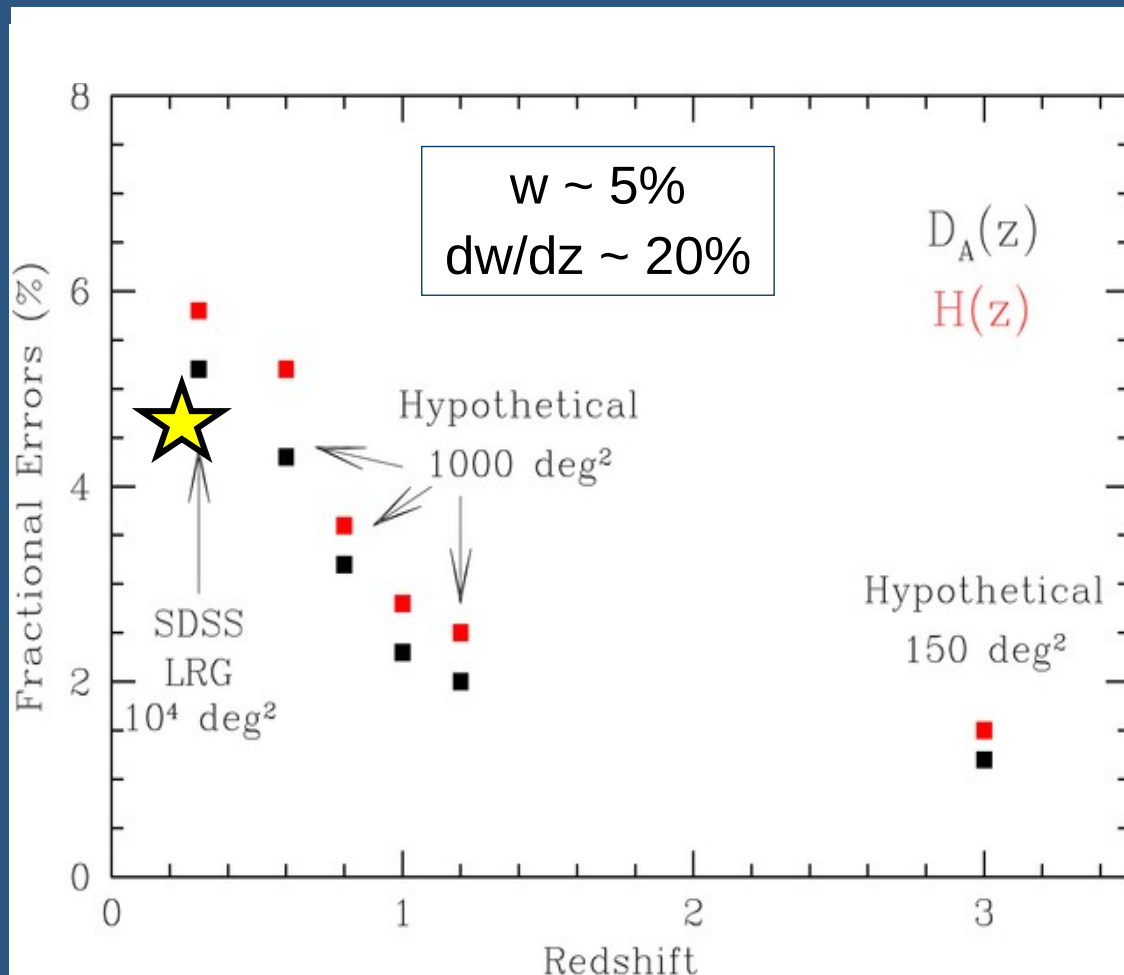
Univ. of Tokyo:

K. Yoshikawa

Y. Suto

Hyper Suprime-Cam and WFMOS on Subaru

3000 thousands of fibers over a 1.5 degree field-of-view
on an 8-meter class telescope (Subaru/Gemini)



Taken from WFMOS Feasibility Study by NOAO, JHU, AAO, Oxford, Durham, Portsmouth, UA

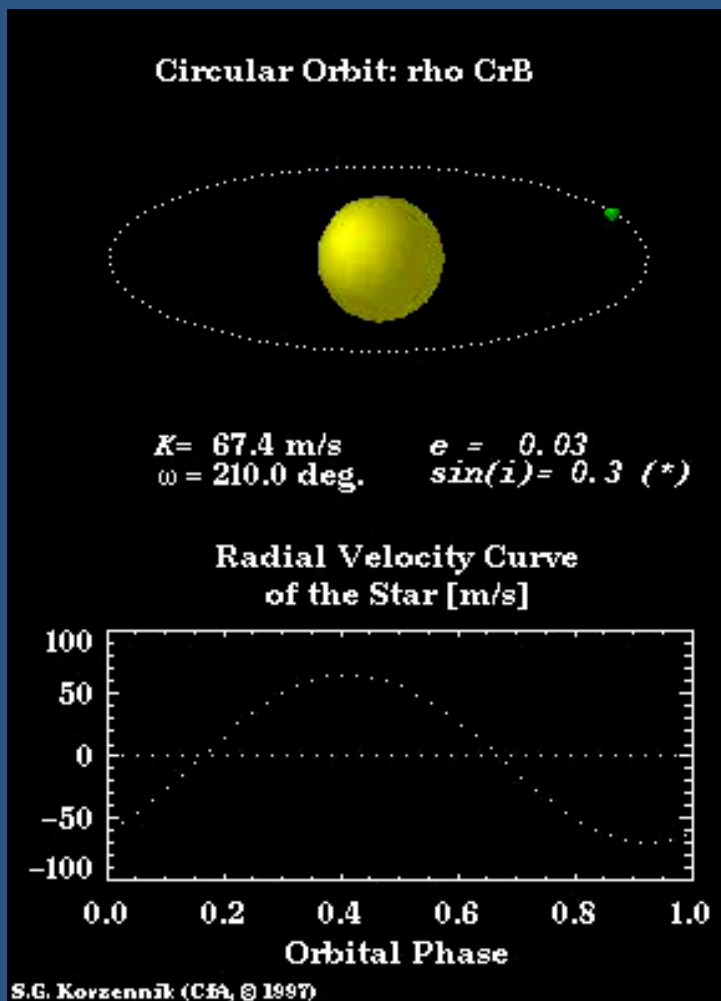
A brief history of the discovery of extrasolar planets

- **1995** : the first extrasolar planet around the main sequence star 51 Pegasi (Mayor & Queloz)
- **1999** : transit of a known planet around HD209458 (Charbonneau et al., Henry et al.)
- **2001** : Na in the atmosphere of HD209458b
- **2003** : first discovery of a planet by transit method *alone* (1.2 day orbital period: OGLE)
- **2005** : spin-orbit misalignment via the Rossiter effect
- 200 extrasolar planets are reported (Sept. 2006)

<http://exoplanets.org/>

Radial velocity of a star perturbed by a planet

Even if planets are not directly observable, their presence can be inferred dynamically



**velocity modulation
of the Sun:**

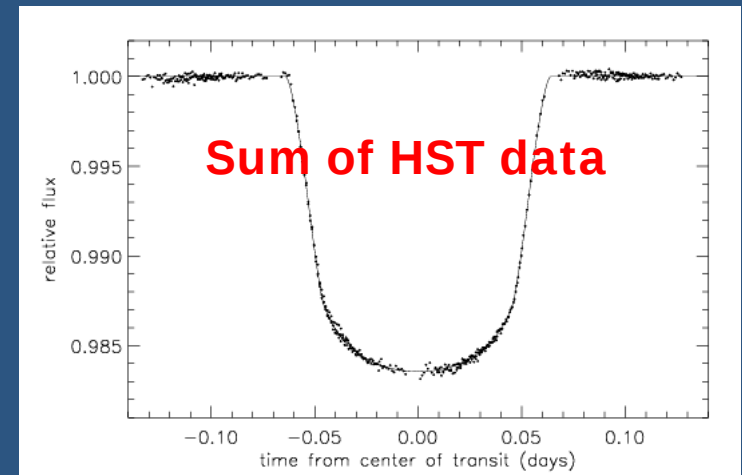
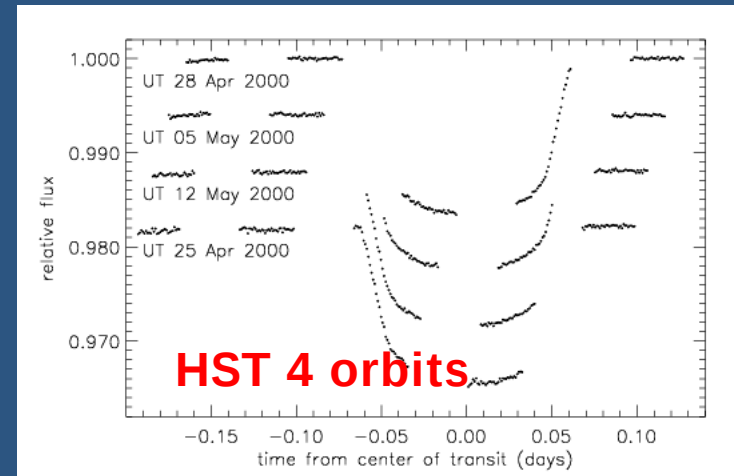
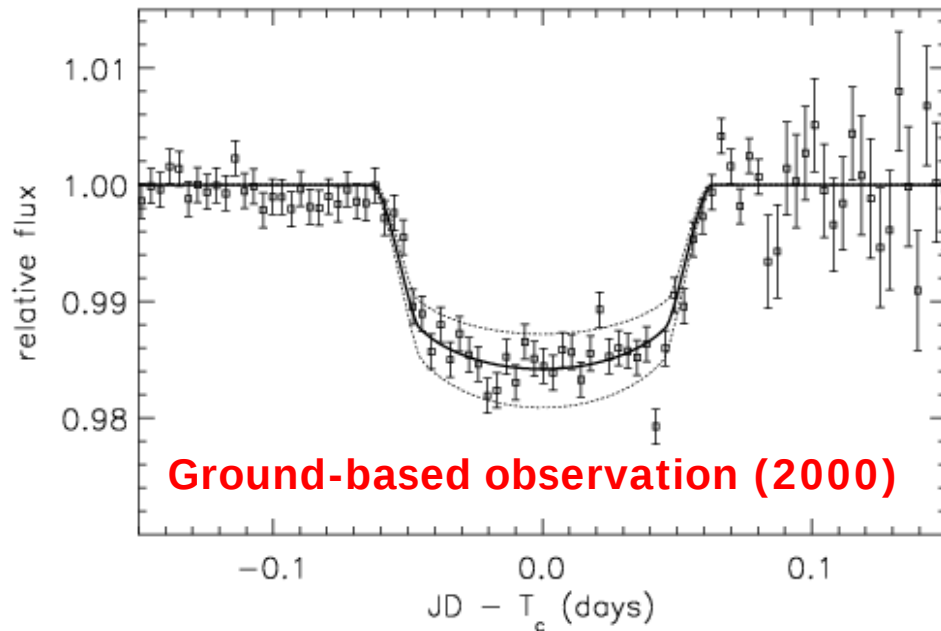
12.5 m/s (Jupiter)

0.1 m/s (Earth)

an accuracy of 1m/s achieved
from the ground observation
 $\Rightarrow$ the current major method in
search for Jupiter-sized planets

the first discovery of the transit of a planet: HD209458

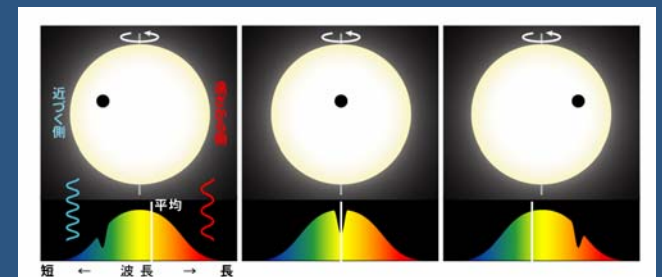
- detected the light curve change at the phase consistent with the radial velocity (Charbonneau et al. 2000, Henry et al. 2000)



Brown et al. (2001)

Extrasolar planet projects at Univ. of Tokyo

- **Search for the planetary atmosphere with Subaru**
 - the most stringent upper limits from ground-based obs.
 - Winn et al. PASJ 56(2004) 655 (astro-ph/0404469)
 - Narita et al. PASJ 57(2005) 471 (astro-ph/0504450)
- **Constraining the stellar spin and the planetary orbital axes from the Rossiter-McLaughlin effect**
 - New analytic formulae (Ohta, Taruya & Suto 2005, ApJ, 622, 1118)
 - First detection (Winn et al. 2005 ApJ, 631, 1215)
- **Search for reflected light from planets**
 - collaboration with Andrew Cameron (St. Andrews Univ.) & Chris Leigh (Liverpool John Moores Univ.)



Subaru observation with HDS



**“Spectro-photometric search
for scattered light from HD209458b”**

S02B-16 on October 24 and 26, 2002

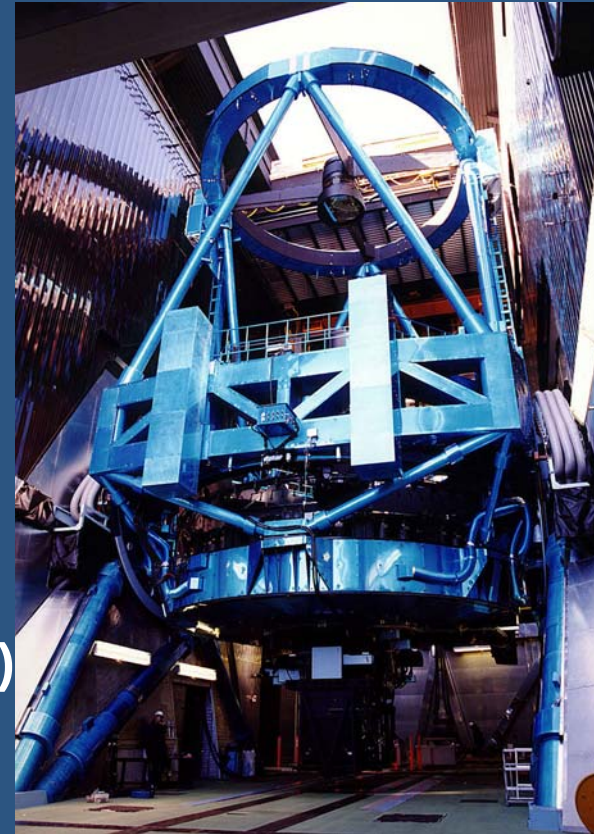
Yasushi Suto, Norio Narita (Univ. of Tokyo)

Toru Yamada, Wako Aoki (National Ast. Obs. Japan)

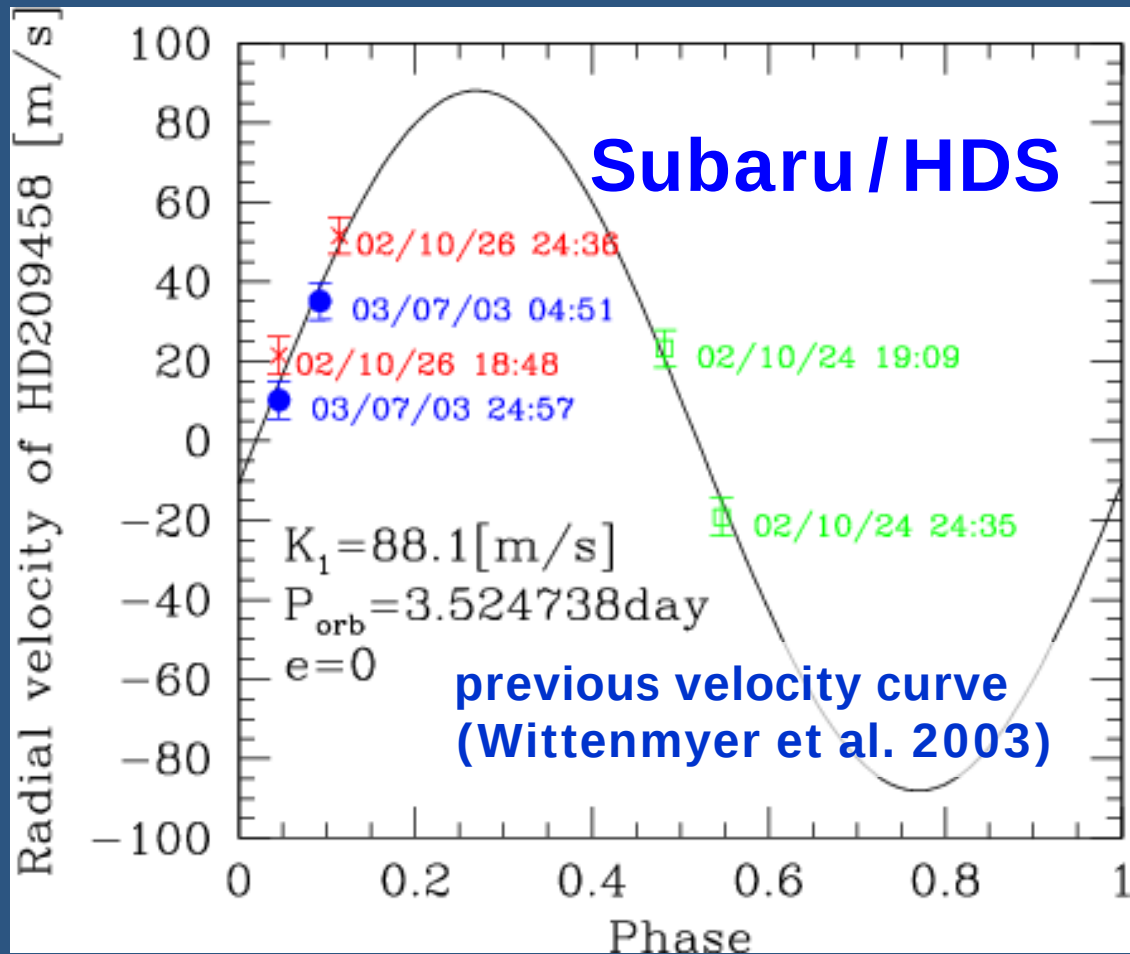
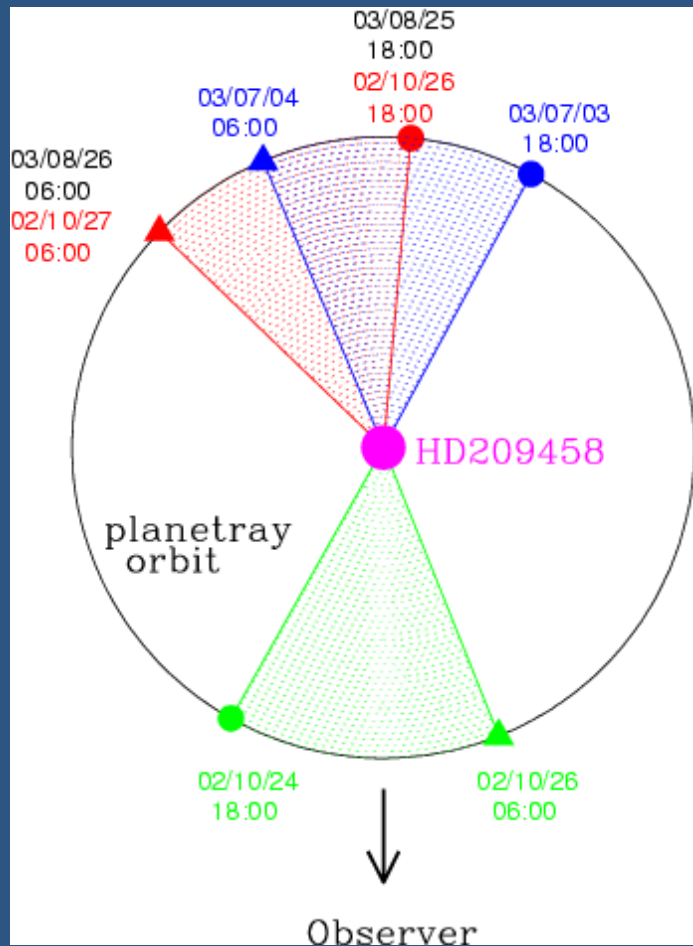
Bun-ei Sato (Kobe Univ.)

Edwin L. Turner (Princeton Univ.)

Josh Winn (Harvard Univ.)



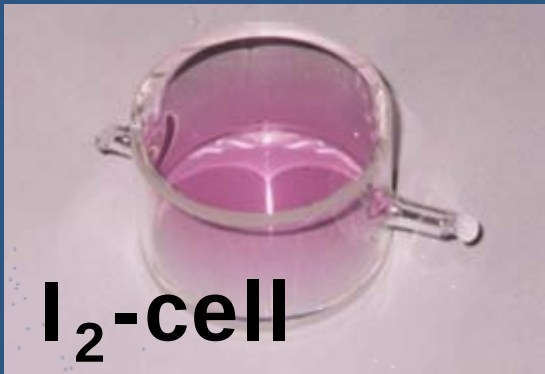
Orbital phase and radial velocity of HD209458b at our observing runs



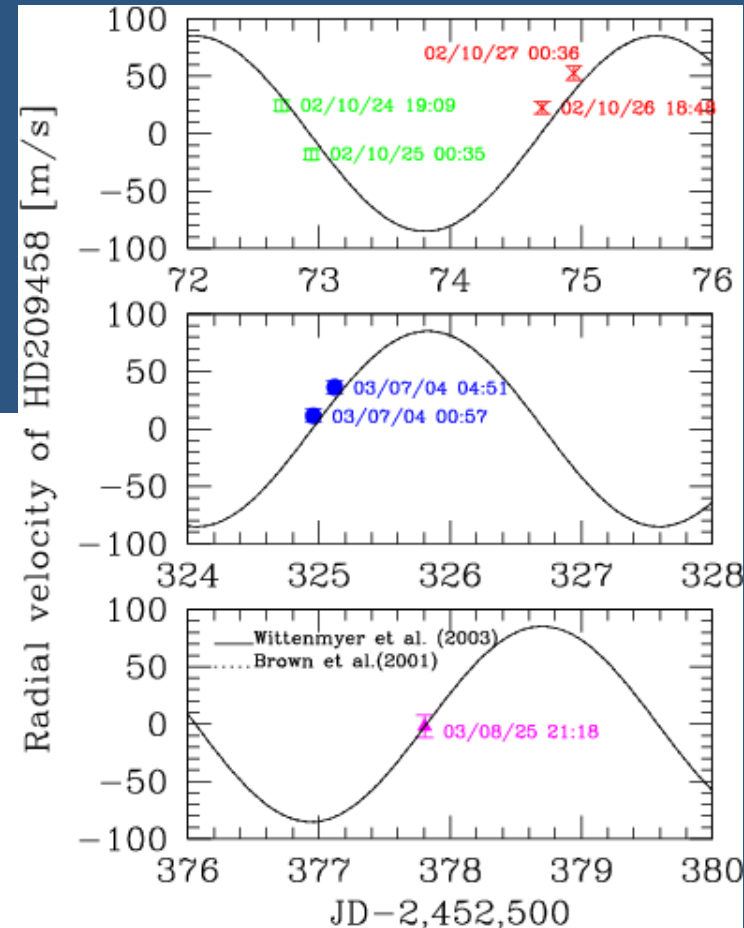
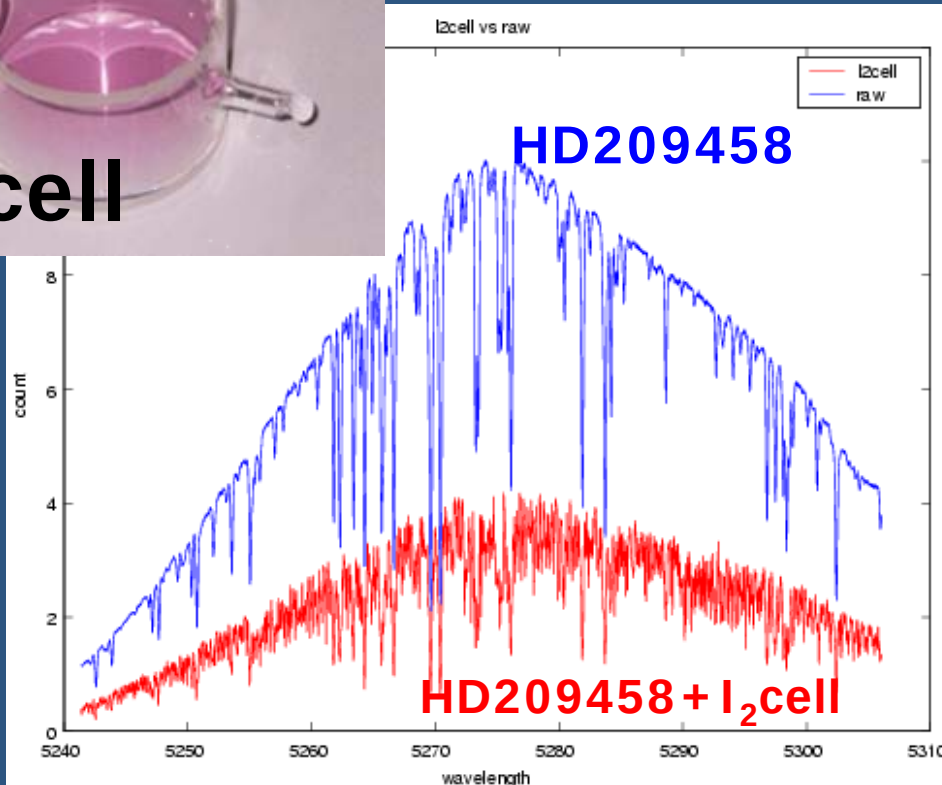
Winn et al. PASJ 56(2004) 655, astro-ph/0404469
Narita et al. PASJ 57(2005)471, astro-ph/0504450

radial velocity measurement with I₂-cell

- Absorption lines of I₂ as accurate standard rulers
- Accuracy $\sim 3\text{m/s}$ achieved

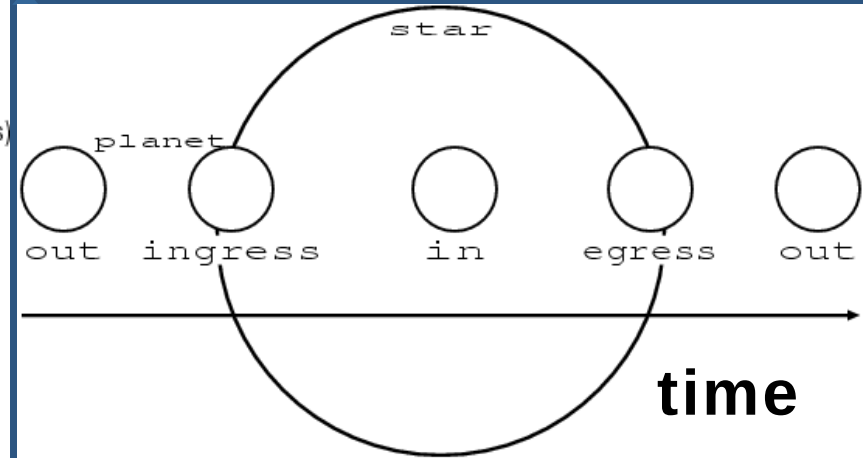
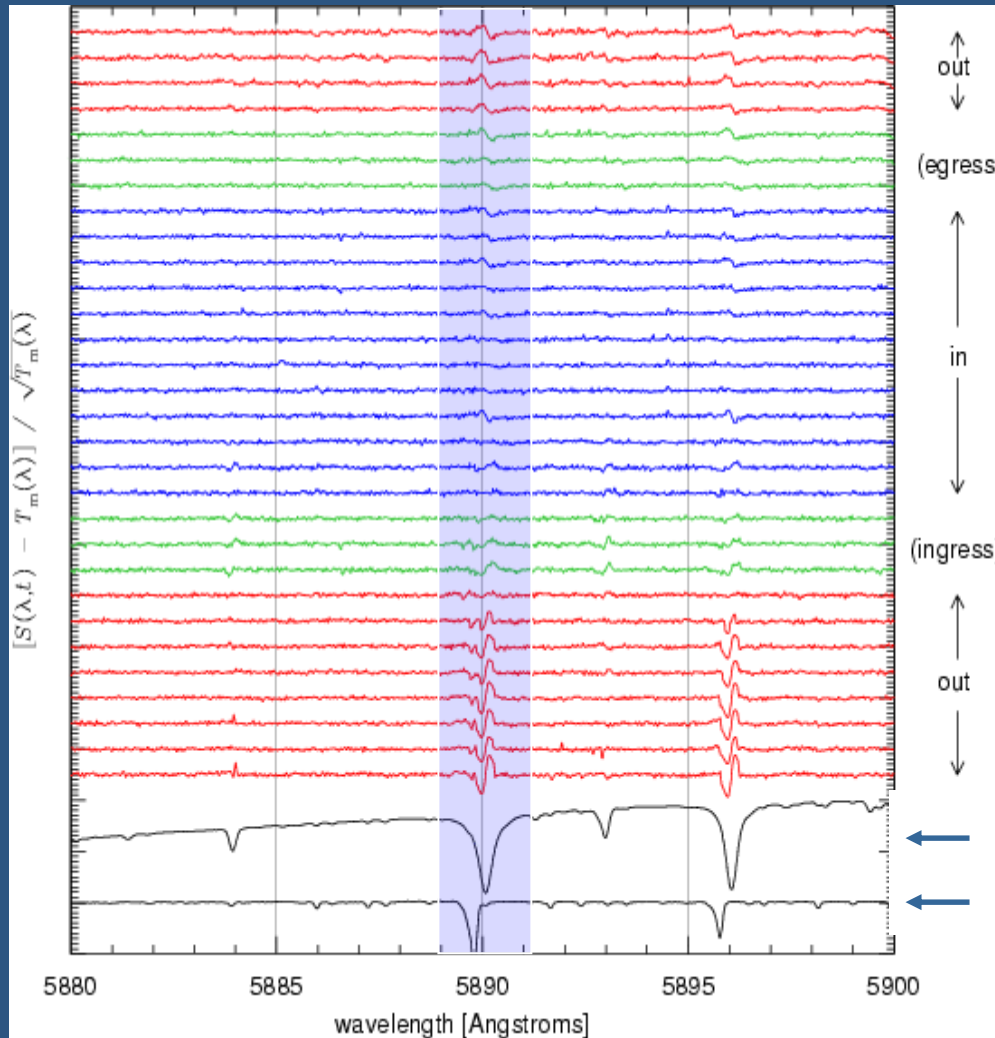


I₂-cell



HDS on Subaru
Winn et al. (2004)

Transit transmission spectroscopy



planetary orbital phase

template

telluric

Narita et al. (2005)

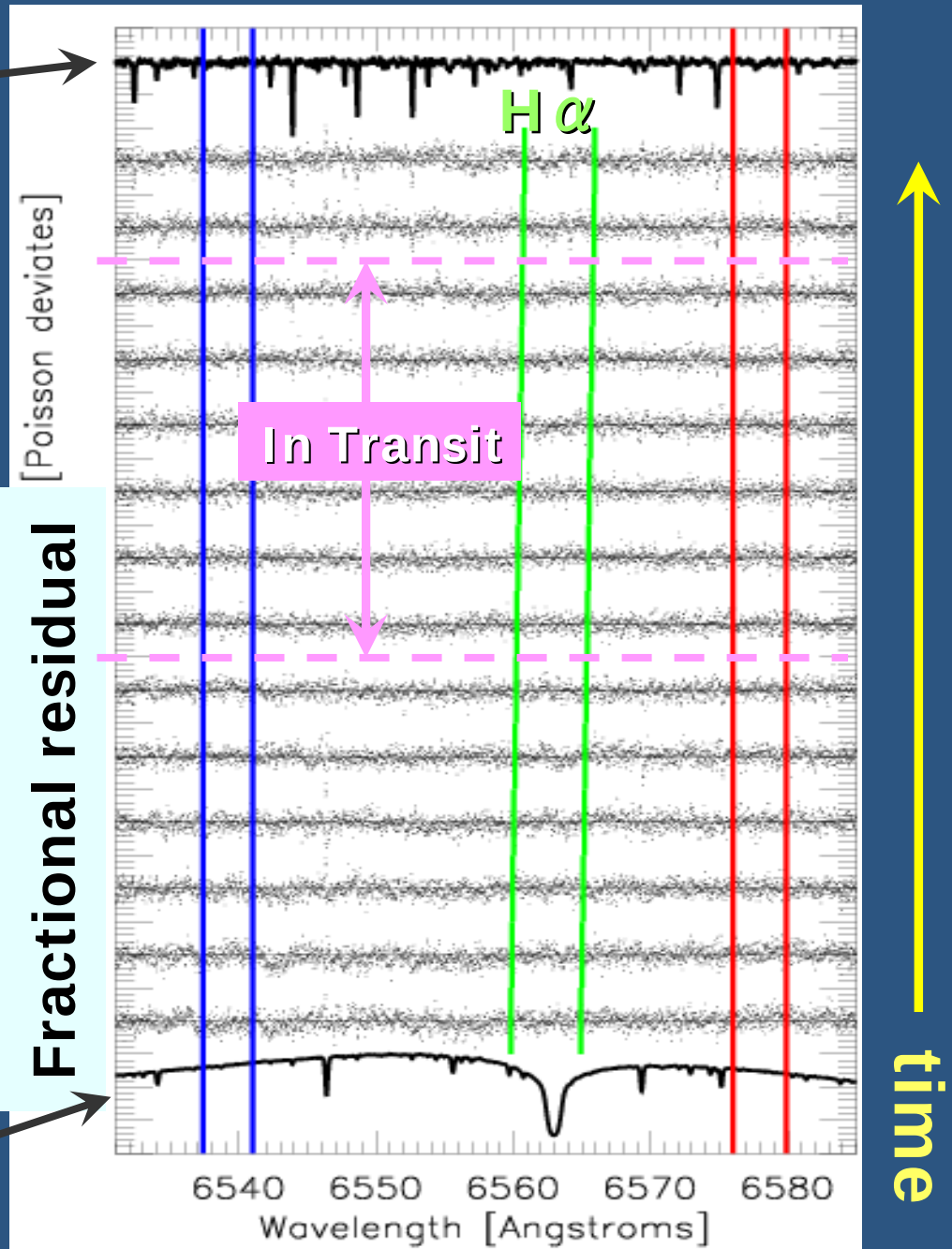
Telluric spectrum

Search for $H\alpha$ absorption due to the atmosphere of HD209458b

Na I (D2)	5889.97 Å
Na I (D1)	5895.94 Å
$H\alpha$	6562.81 Å
$H\beta$	4861.34 Å
$H\gamma$	4340.48 Å

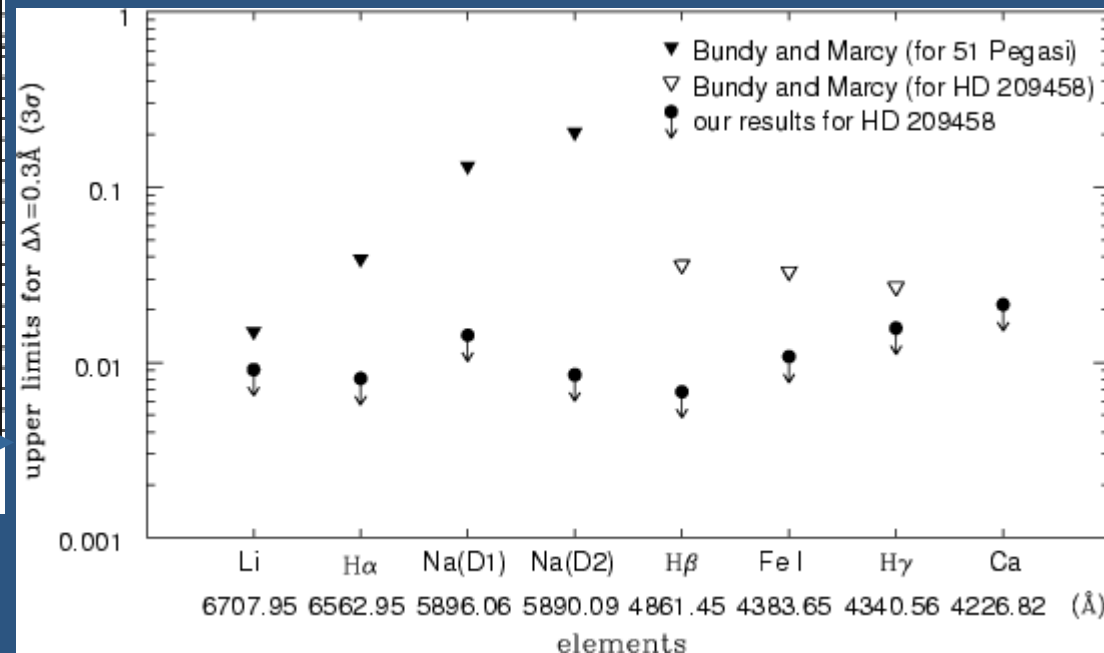
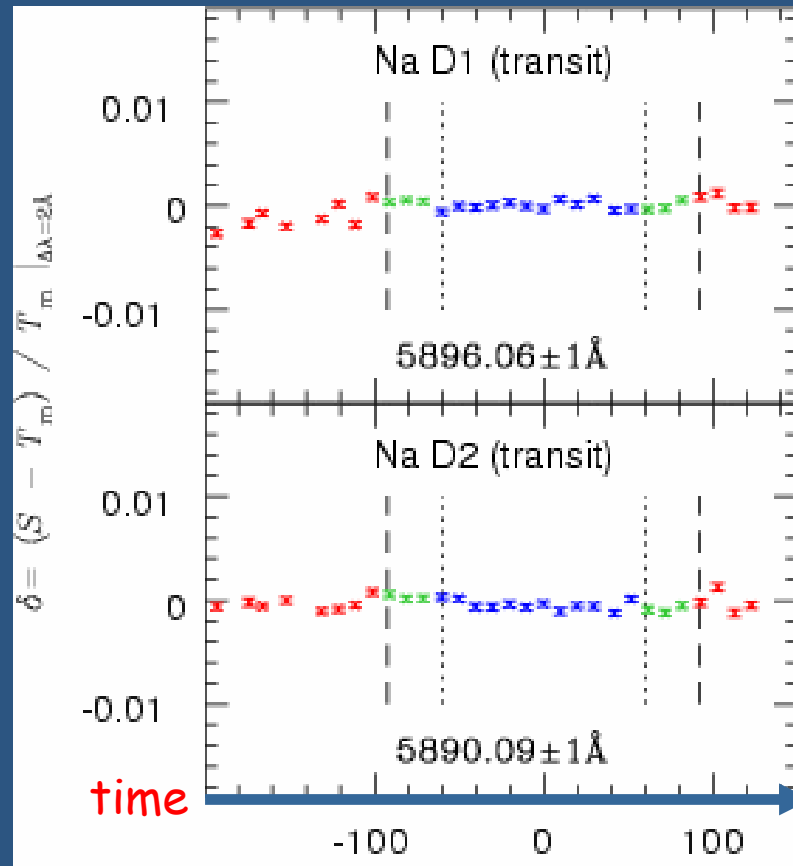
Template stellar spectrum

Winn et al. (2004)



most stringent upper limits from ground-based optical observations

Comparison with previous results
for 0.3 angstrom bandwidth
(Bundy and Marcy 2000)



Narita et al. (2005)

Approaching part

Receding part

planet

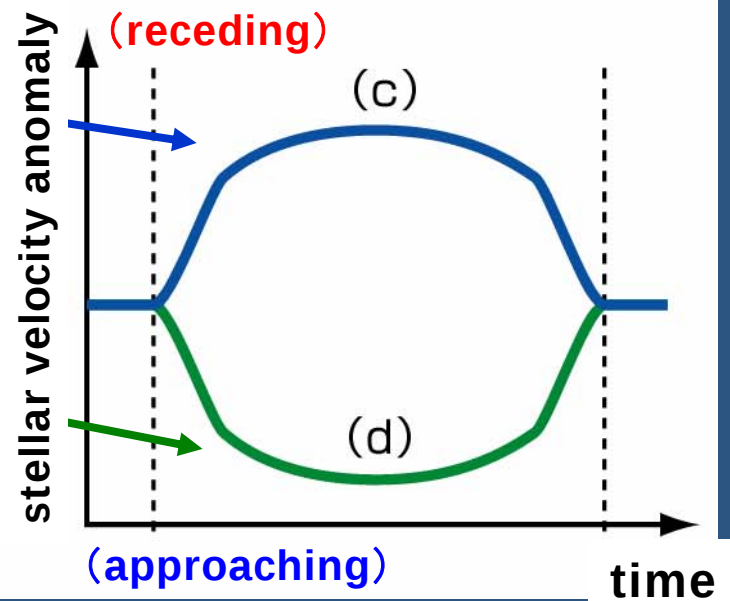
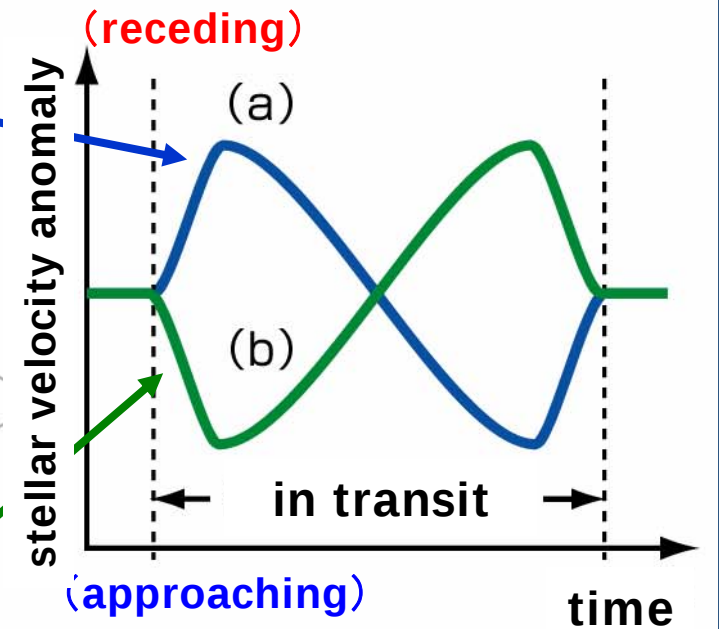
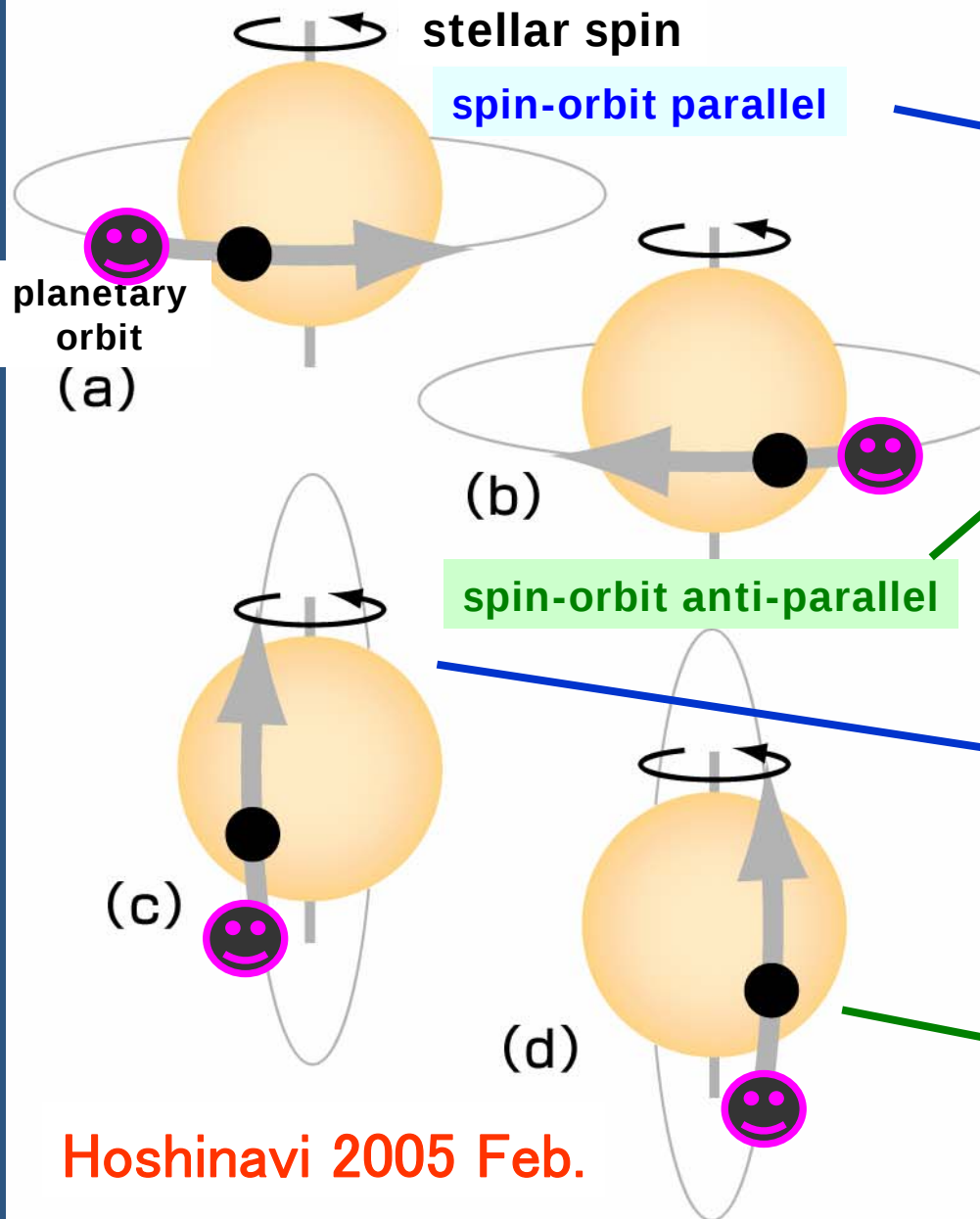
star

Shadow of the planet

Wavelength shift of line center

- Time-dependent asymmetry in the stellar Doppler broadened line profile
 - an apparent anomaly of the stellar radial velocity
- originally discussed in eclipsing binary systems
 - Rossiter (1924)
 - McLaughlin (1924)

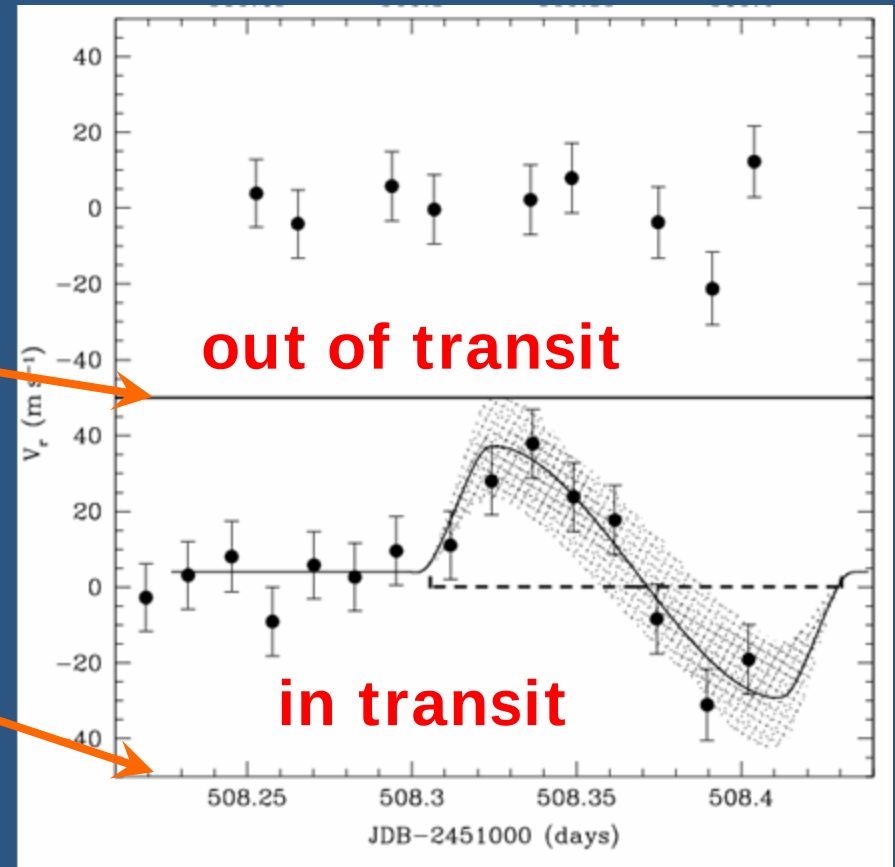
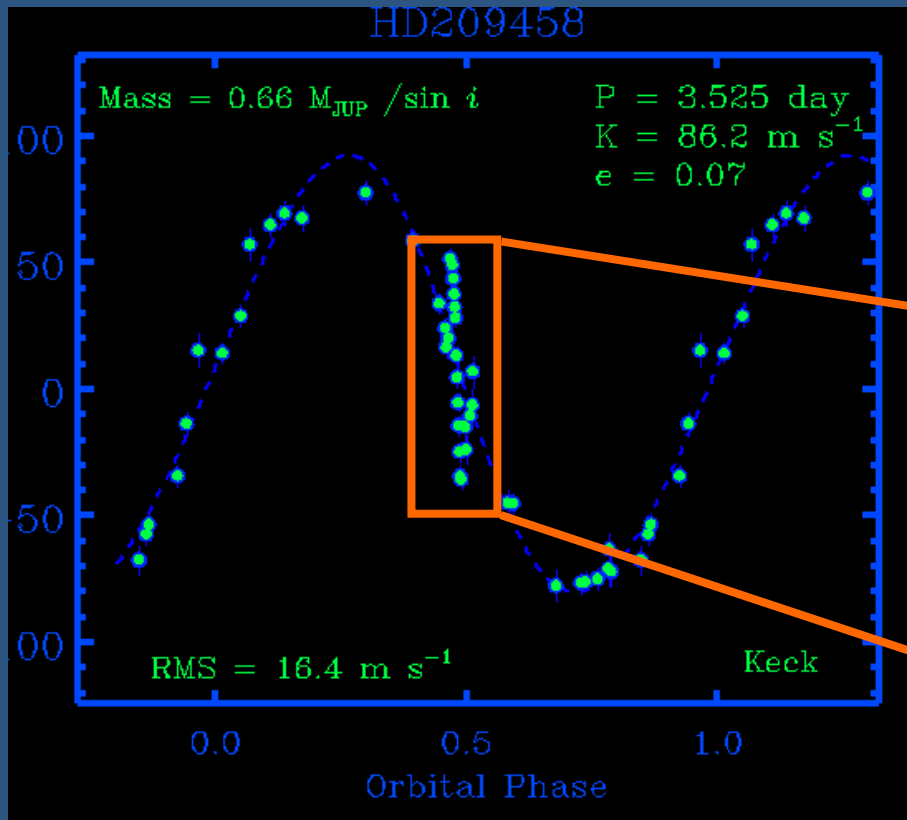
Velocity anomaly due to the Rossiter effect



Hoshinavi 2005 Feb.

Previous result of the Rossiter-McLaughlin effect for an extrasolar transit planetary system HD209458

Origin of angular momentum



HD209458 radial velocity data <http://exoplanets.org/>

Stellar rotation and planetary orbit
Queloz et al. (2000) A&A 359, L13
ELODIE on 193cm telescope

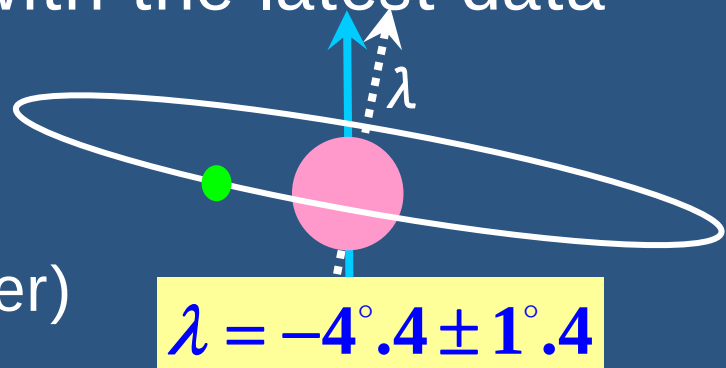
Measurement of Spin-Orbit alignment in an Extrasolar Planetary System

- **Joshua N. Winn (MIT)**, R.W. Noyes, M.J. Holman, D.B. Charbonneau, Y. Ohta, A. Taruya, Y. Suto, N. Narita, E.L. Turner, J.A. Johnson, G.W. Marcy, R.P. Butler, & S.S. Vogt
 - **ApJ 631(2005)1215 (astro-ph/0504555)**



Precision analysis of the Rossiter-McLaughlin effect for HD209458

- perturbation formula by Ohta et al. (2005)
- HD209458 re-examined with the latest data
 - radial velocity data (Keck)
 - optical photometry (HST)
 - infrared photometry (Spitzer)
- **the first detection of the misalignment between the stellar spin and the planetary orbital axes by (-4.4 ± 1.4) deg**
 - more than an order-of-magnitude improvement of the previous error-bar (maybe useless but impressive result !)
 - c.f., 6 degree misalignment for the Solar system



HD209458 parameter fit

$$\chi^2 = \sum_{n=1}^{83} \left(\frac{v_{\text{obs}} - v_{\text{model}}}{\sigma_v} \right)^2 + \sum_{n=1}^{417} \left(\frac{f_{\text{obs}} - f_{\text{model}}}{\sigma_f} \right)^2 + \left(\frac{t_{2\text{nd,obs}} - t_{2\text{nd,model}}}{\sigma_t} \right)^2$$

12 free parameters

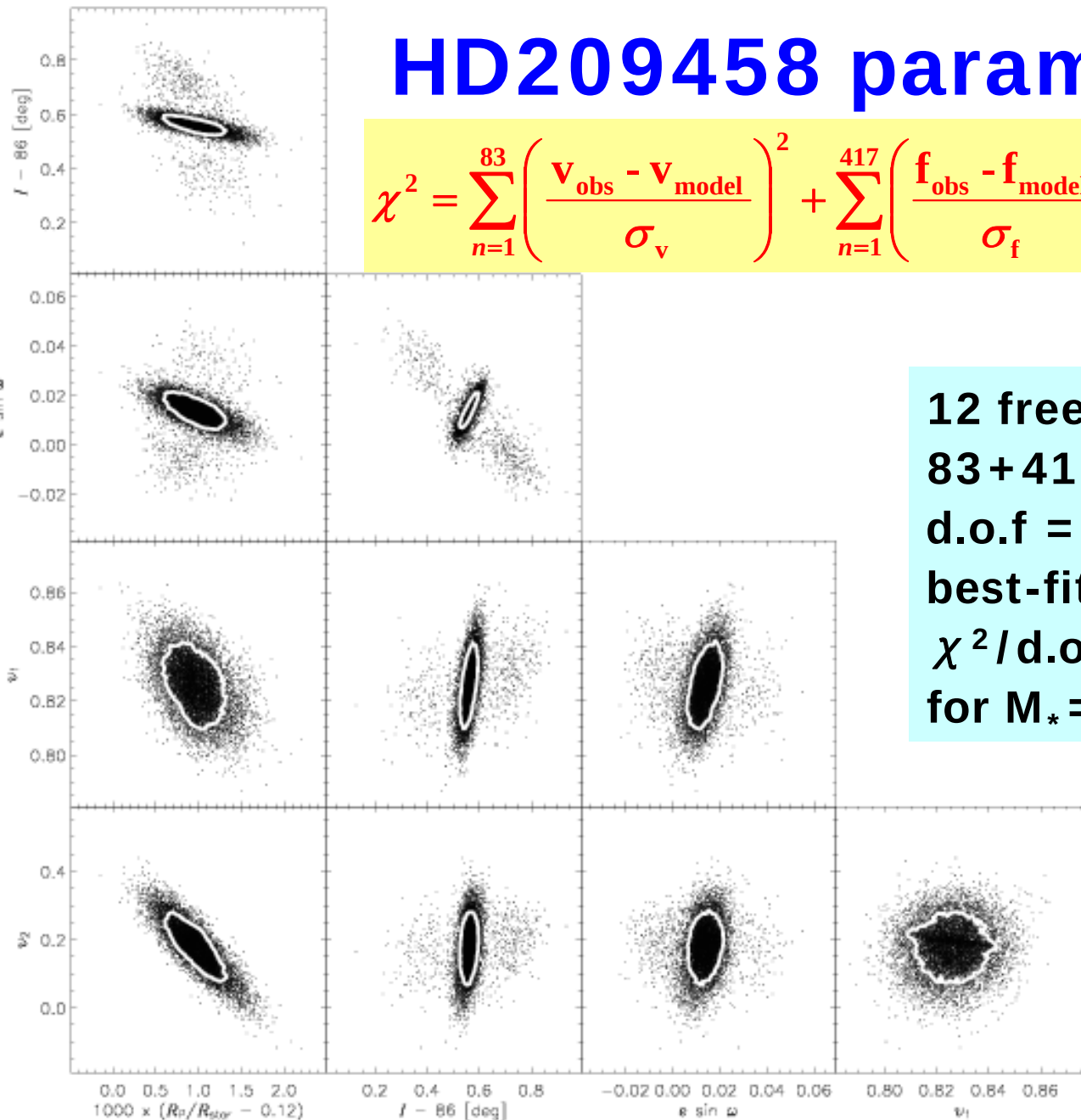
83+417+1 data points

d.o.f = 83+417+1-12=489

best-fit :

$\chi^2/\text{d.o.f} = 528/489 = 1.08$

for $M_* = 1.06 M_{\text{sun}}$

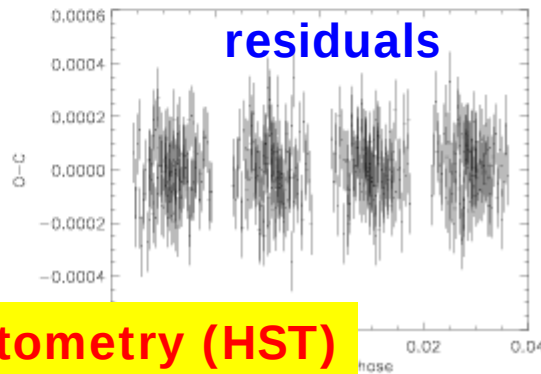
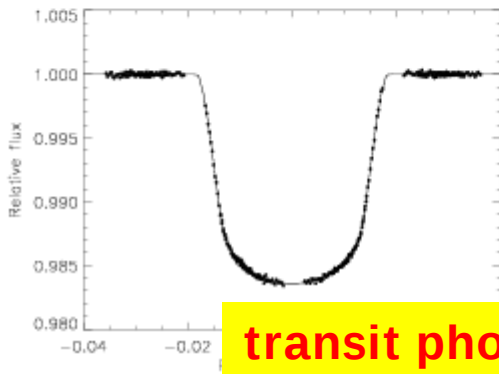
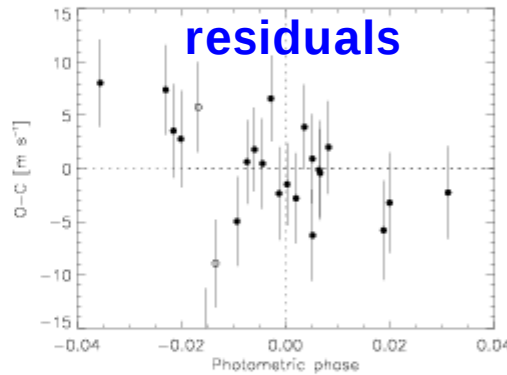
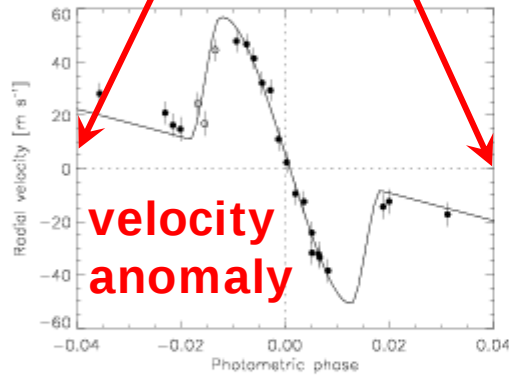
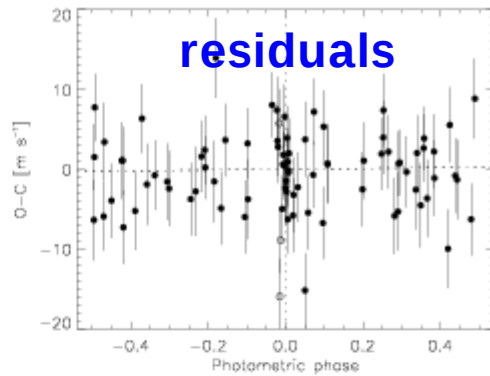
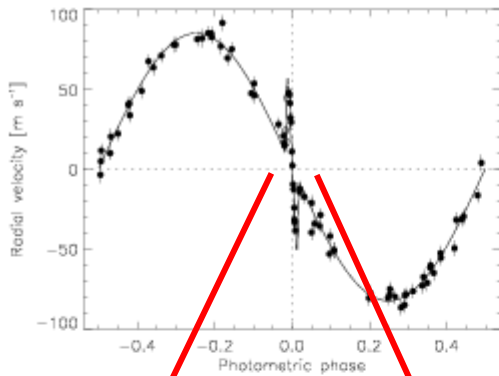


Winn et al.

astro-ph/0504555

ApJ 631(2005)1215

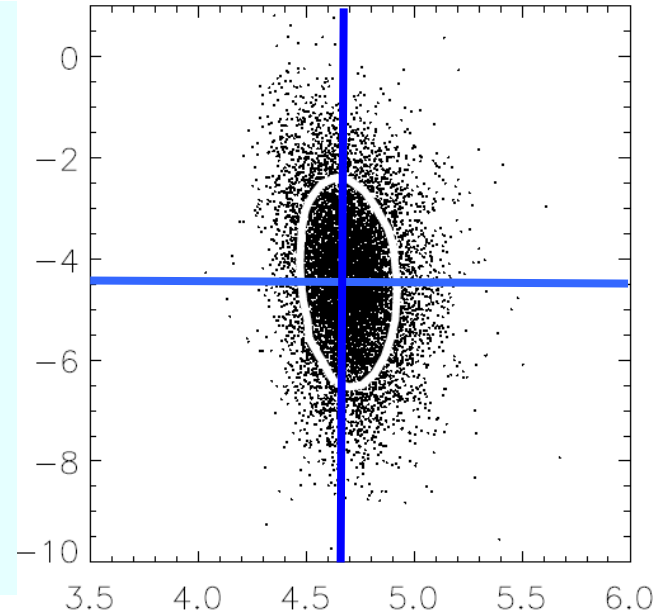
radial velocity (Keck)



transit photometry (HST)

first detection
of non-zero λ !

misalignment angle [deg]



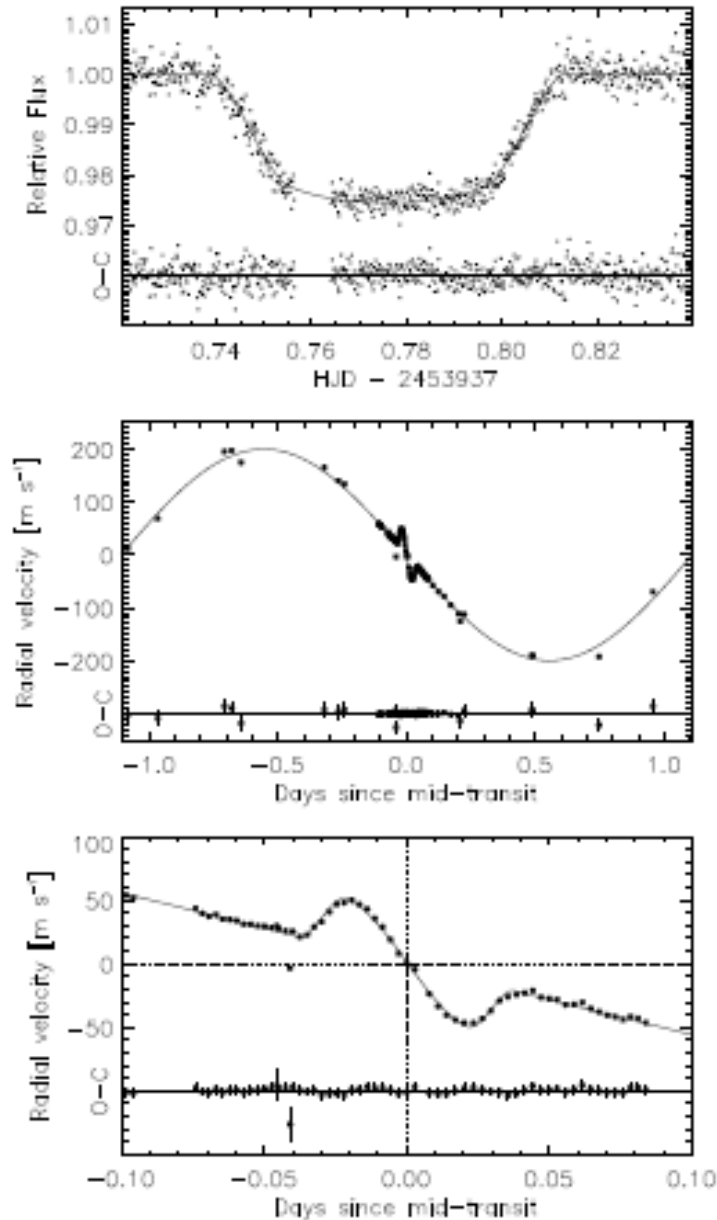
(projected) stellar spin velocity [km/s]

$$\lambda = -4^{\circ}.4 \pm 1^{\circ}.4$$

3σ detection !

Winn et al. astro-ph/0504555 ApJ 631(2005)1215

another example: HD189733



■ HD 189733

- Bouchy et al. (2005)

- K-dwarf
(V=7.7)+1.15M_Jplanet

- 2.2 day orbital period

- 2.5% photometric transit signal

- Winn et al. ApJL, submitted (astro-ph/0609506)

- $\lambda = -1.4^\circ \pm 1.1^\circ$

- consistent with zero

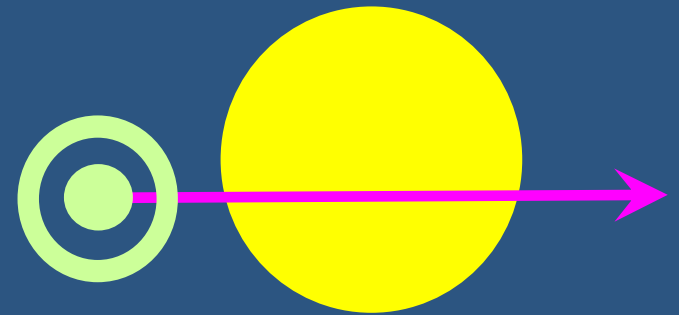
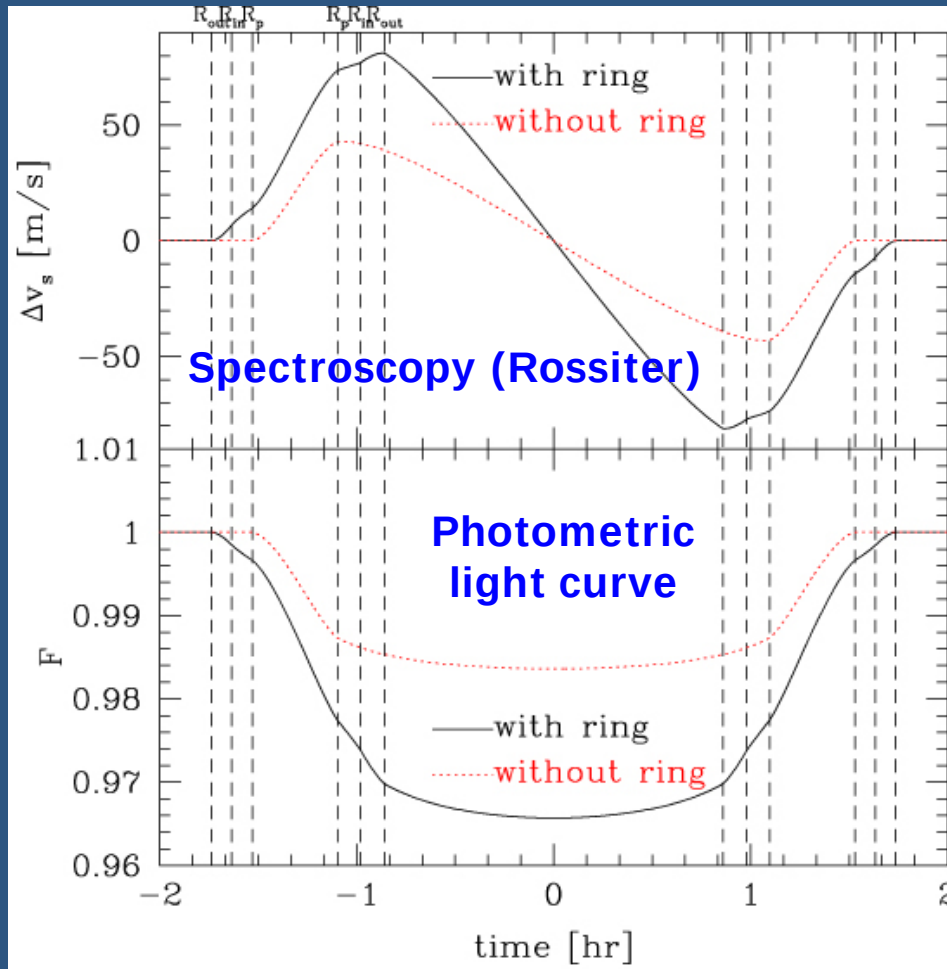
Discovery of the Uranus rings



■ Serendipitous discovery

- Uranus's transit against a background star (Elliot et al. 1977)
- Neptune's ring was discovered also by transit technique in 1986
- **Transit proved to be useful in detecting rings of the Solar planets !**

Signatures of planetary rings



- Ring's inner and outer radii, gap, planet's radius imprints strong features in the photometric and spectroscopic data
- Statistical analysis of the residuals with respect to the best-fit ringless model

Ohta, Taruya & Suto, in preparation

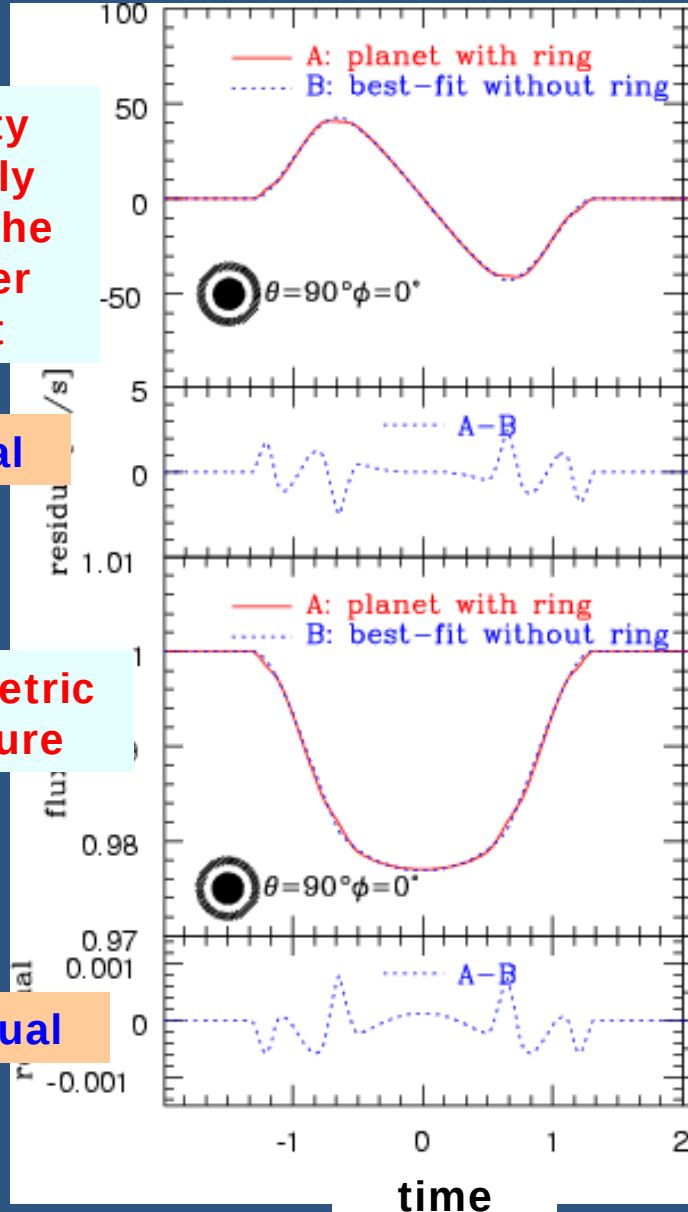
Detectability of a ring

Velocity
anomaly
due to the
Rossiter
effect

residual

Photometric
signature

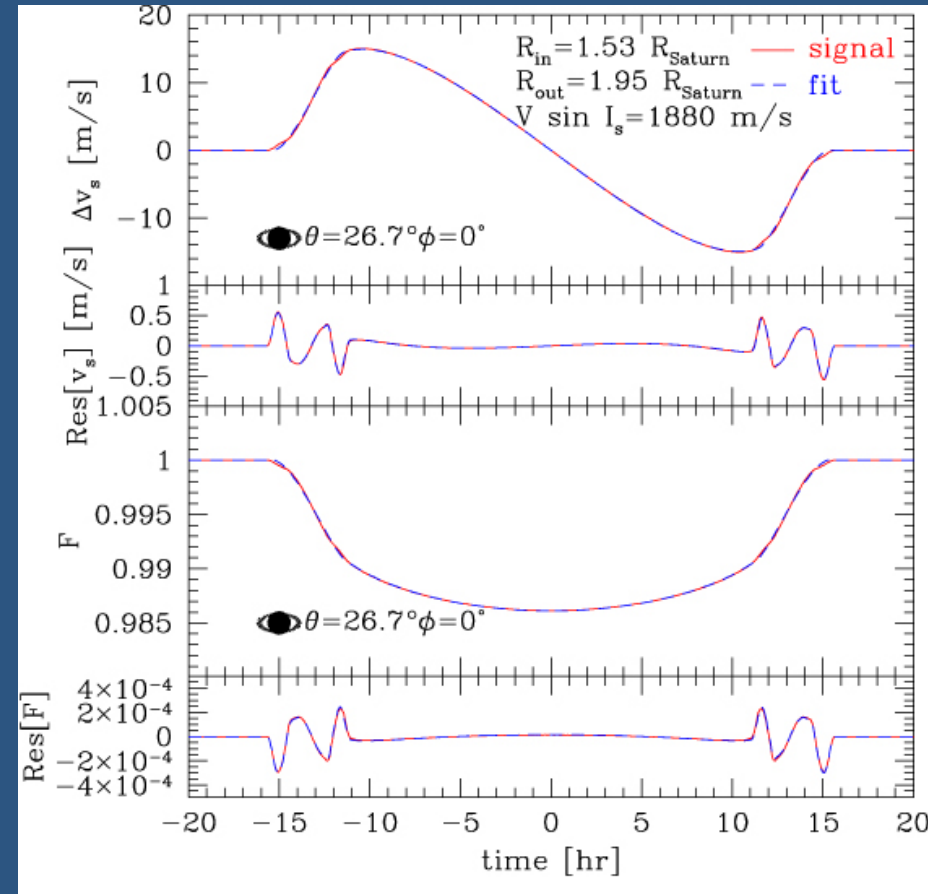
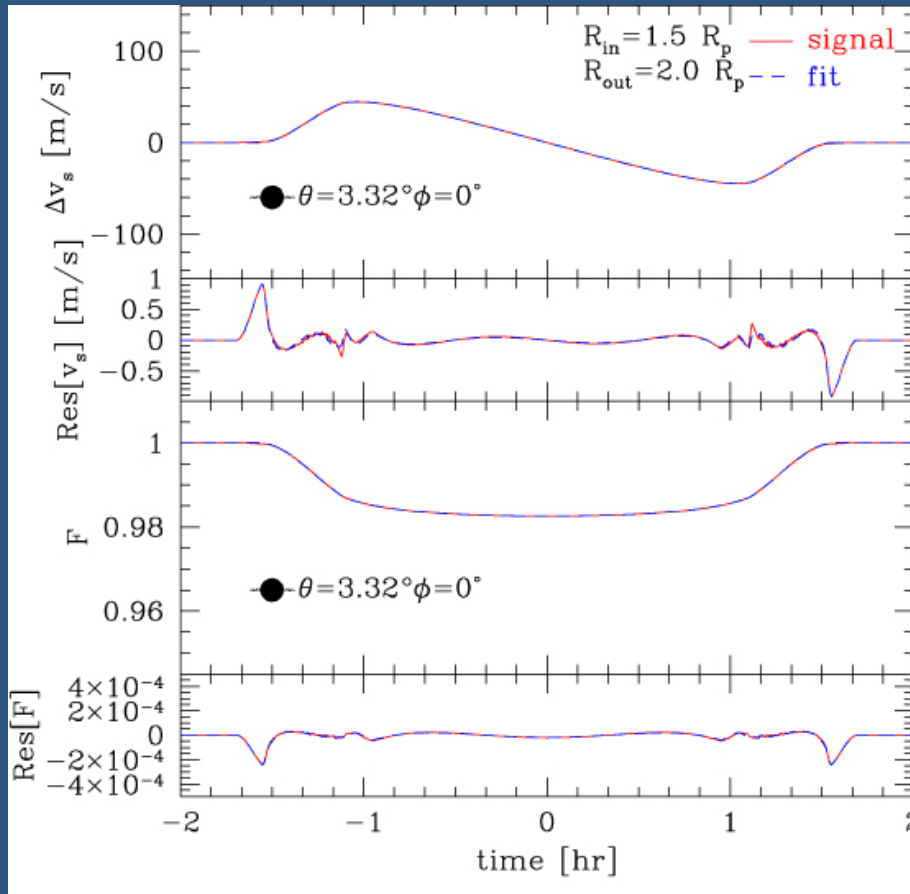
residual



- a hypothetical ring around HD209458
 - $1.5R_{\text{pl}} < R_{\text{ring}} < 2R_{\text{pl}}$
 - deviation from a best-fit single planet
 - $\delta v \sim 1 \text{ m/s}$
 - $\delta F / F \sim 0.1 \%$
- marginally detectable level even with the current technology

Ohta, Taruya & Suto, in preparation

How about hot Jupiter and Saturn rings ?



Ohta, Taruya & Suto, in preparation

- Hot Jupiter: edge-on rotation due to the tidal locking
- Saturn: 30 deg. inclined, but spin of the Sun is small
- Worse in either case, but still detectable potentially ($S/N=1$)

A possible roadmap of sciences of extrasolar planet

- Discovery phase of gas giant planets (1995-)
 - Discovery of planetary atmosphere (2002)
 - Detailed spectroscopic study of planets
 - Discovery of IR emission of planets (2005)
-
- 
- Discovery of visible reflection light
 - Discovery of planetary ring
 - Discovery of extrasolar planetary satellite
 - Discovery of terrestrial planets
-
- ***Discovery of Habitable planet***
 - ***Identifying Biomarker***
 - Red-edge of extrasolar plant ?
 - ***Discovery of Extraterrestrial life***

Astrobiology ? Not yet

- Discovery of extrasolar planets is a wonderful breakthrough in astronomy (and philosophy, maybe)
- But mere discovery has no biological information
- **How can we identify the signature of life ?**

- **Biomarker**

- Suppose our earth is  located at 10pc away. Can we identify any signature of life from photometric and spectroscopic data alone ?

- **Earth-shine**

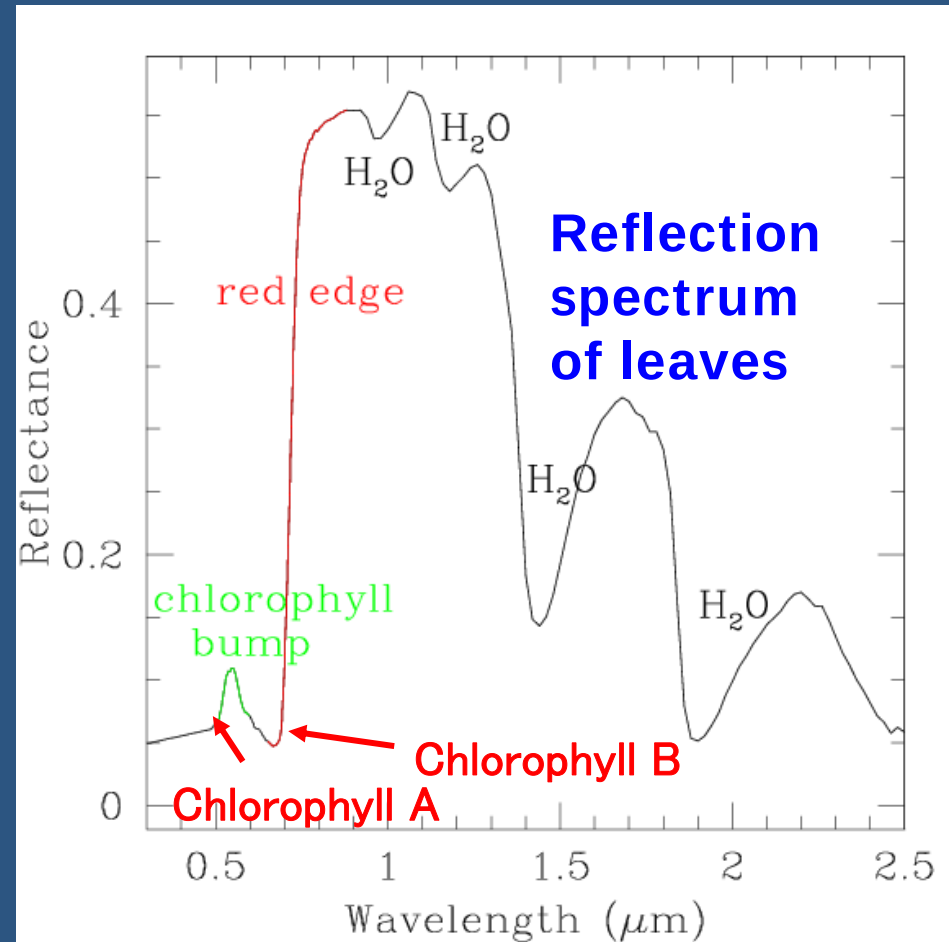


<http://modarch.gsfc.nasa.gov/>

<http://www.nasa.gov/home/index.html>

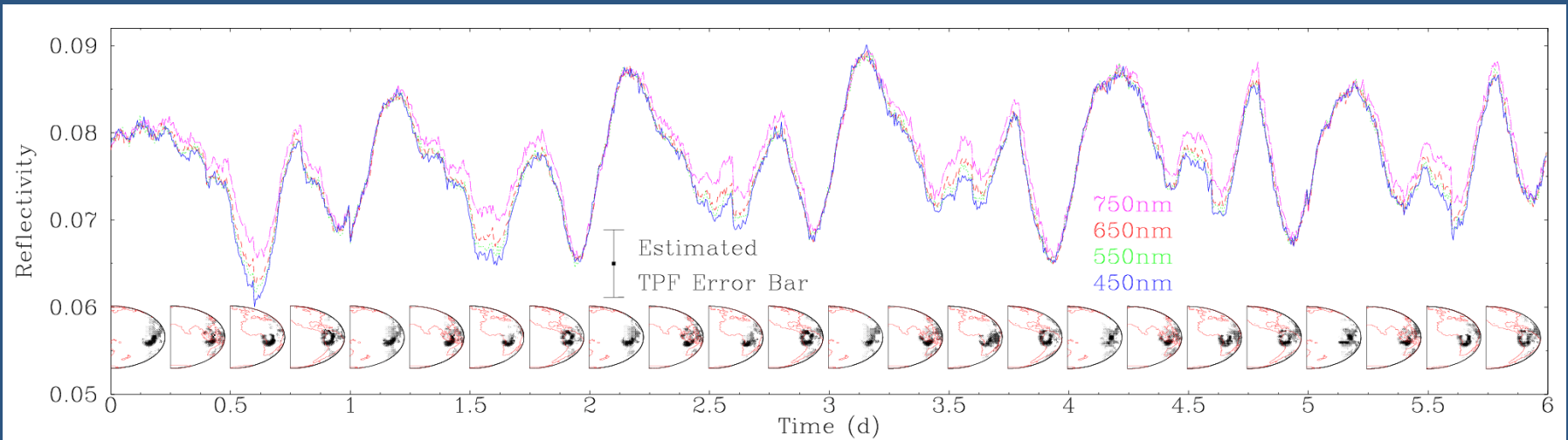
Red edge of (*extrasolar*) plants: a biomarker in *extrasolar planets*

- Significant reflectivity of leaves of terrestrial planets for $\lambda > 7000 \text{ \AA}$
- An interesting (maybe unique) candidate for a biomarker ?
- *extrasolar plants* as a biomarker in *extrasolar planets*



Seager, Ford & Turner
astro-ph/0210277

Expected daily change of the reflected light from the earth

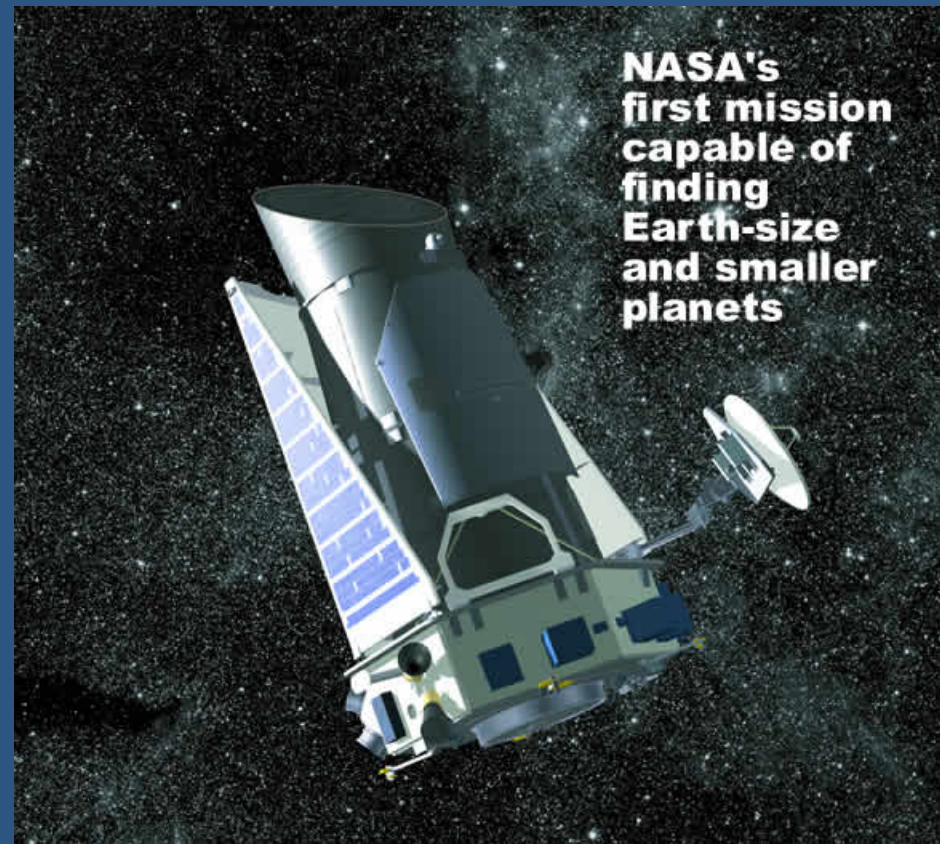
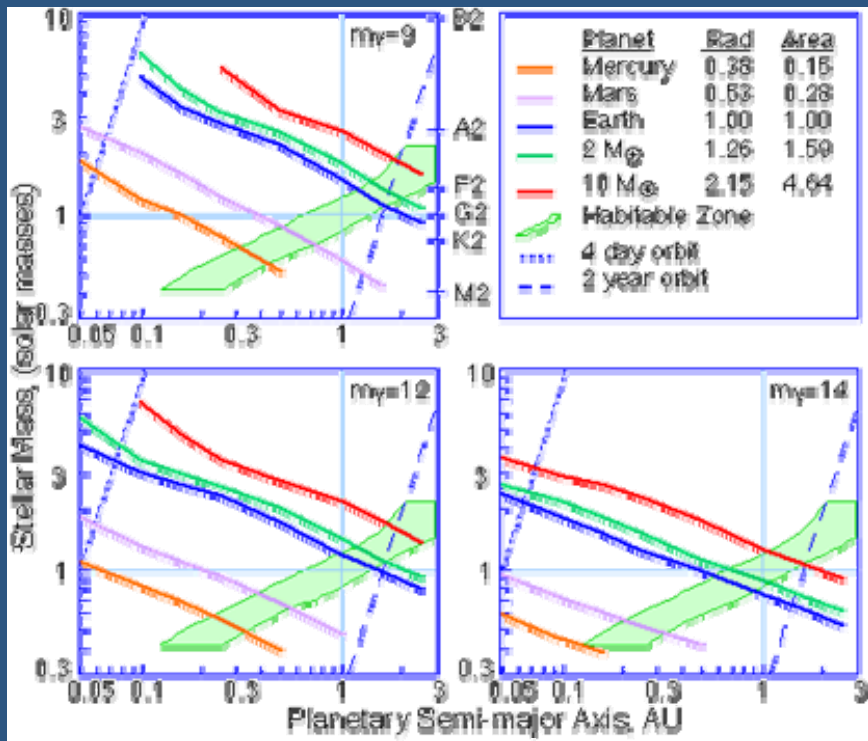


Ford, Seager & Turner: Nature 412 (2001) 885

- **Assume** that the earth's reflected light is completely separated from the Sun's flux !
 - TPF (Terrestrial Planet Finder) in (10~20) years from now ?
- **Periodic change of 10 % level** due to different reflectivity of land, ocean, forest, and so on
- Cloud is the most uncertain factor: **weather forecast !**

Kepler mission (June 2008 launch?)

differential photometry survey of transit planets
expect to discover > 50 terrestrial planets in 4 years ?



<http://kepler.nasa.gov/>

Future of transiting planet research: follow Hantaro Nagaoka

■ Nagaoka's Saturn model of atom

- Nagaoka: Phil. Mag. 7(1904) 445
- $\Rightarrow$ quantum atomic physics

■ Transit planets

- Orbital angular momentum (**L**):
radial velocity

- Spin of star (**S**): Rossiter effect

- Spin of planet (**s**): ring, satellite

■ From planets to atomic physics

■ From atomic physics to planets

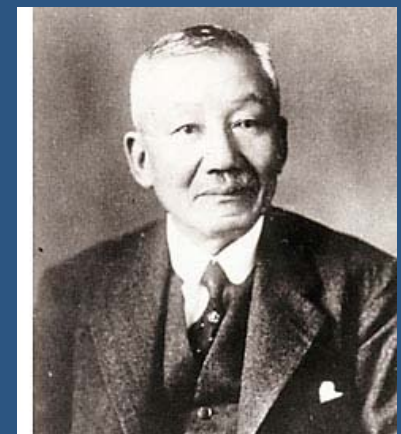


写真 2: 長岡半太郎 (1865-1950)

Extra-solar planet can have a ring ?

■ Final sec. of the Nagaoka paper

- There are various problems which will possibly be capable of being attacked on the hypothesis of a Saturnian system
- The rough calculation and rather unpolished exposition of various phenomena above sketched may serve as a hint to a more complete solution of atomic structure



写真 2：長岡半太郎（1865－1950）