Colors of a second Earth: towards exoplanetary remote-sensing



正式大学



band = 1,2,3,4,5 $\tau = \tau_0$

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Department of Physics, the University of Tokyo & Global Scholar, Department of Astrophysical Sciences, Princeton University seminar at Centre de Recherche Astrophysique de Lyon 14:00 October 20, 2011@ENS-Big CECAM

Nightfall: We didn't know anything



no "night" except the total eclipse due to another planet every 2049 years on a planet "Lagash"

People realized the true world for the first time through the darkness full of "stars" (Issac Asimov: Nightfall)

History of exoplanet discovery

Number of planets by year of discovery



exoplanet.eu (10/10/11)

exoplanet projects in my group at the Univ. of Tokyo

- Constraining the stellar spin and the planetary orbital axes from the Rossiter-McLaughlin effect
 - analytic perturbation formulae (Ohta et al. 2005, ApJ, 622, 1118; Hirano et al. 2010, ApJ, 709, 458; 2011 ApJ, in press)
 - First accurate detection (Winn et al. 2005 ApJ, 631, 1215)
 - application to ring detection (Ohta et al. 2009, ApJ, 690, 1)
- Colors of a second earth
 - Estimating the fractional areas of surface components from simulated photometry data (Fujii et al. 2010 ApJ, 715, 866;2011 ApJ, 738, 184)

What we have learned so far...

- Planets are not rare, but fairly common
 - >10 percent of sun-like stars have planets
- Diversity of planetary systems
 - Hot Jupiter, super earth,,,
 - Prograde/retrograde/polar-orbit planet
- Various observational approaches
 - High-dispersion spectroscopy (radial velocity), precise photometry (transit, micro-lens), direct imaging
 - Planetary atmosphere
 - Reflected light from planet

What's next ?

Kepler mission (March 6, 2009 launch) Photometric survey of transiting planets Searching for terrestrial (and habitable) planets







1st public data release 706 transiting planet candidates (Borucki et al. arXiv:1006.2799)

http://kepler.nasa.gov/

O₃: The Occulting Ozone Observatory

O3: The Occulting Ozone Observatory



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The New Worlds Mission: search for terrestrial planets



http://newworlds.colorado.edu/

Visible-band mission with 2-4m apertuer@L2

- Occulter mission @7x10⁴km away
- Photometric and spectroscopic monitor of planets
- Search for biomarker
- US+UK project; Univ. of Colorado

Conventional biomarkers (signature of life)



A-band@0.76 μ m B-band@0.69 μ m H_2O **0.72**, 0.82, 0.94 μ m Chappuis band @(0.5-0.7) μ m

• Hartley band $@(0.2-0.3) \mu$ m

Kasting et al. arXiv:0911.2936 "Exoplanet characterization and the search for life"

Earth's IR spectrum and biomarkers



O₃@9.6 μ m
 Good tracer of O₂
 H₂O
 <8 μ m, >17 μ m
 CH₄@7.7 μ m
 Biotic origin?

Kasting et al. arXiv:0911.2936 "Exoplanet characterization and the search for life"

Red edge of *(extrasolar) plants:* a biomarker in *extrasolar planets*

Red-edge

- Significant increase of reflectivity of leaves on Earth (terrestrial planets) for λ >7000Å
- An interesting and unique biomarker ?
- Widely used in the remote-sensing of our Earth



Seager, Ford & Turner astro-ph/0210277

Vesto Melvin Slipher (1875-1969)

Red-edge as a biomarker (at least) in 1924 !

 Discovered redshifts of "spiral nebulae" now known as galaxies

"Observations of Mars in 1924 made at the Lowell Observatory: II spectrum observations of Mars" PASP 36(1924)261



reflection spectrum. The Martian spectra of the dark regions so far do not give any certain evidence of the typical reflection spectrum of chlorophyl. The amount and types of vegetation required to make the effect noticeable is being investigated by suitable terrestrial exposures. <u>Astrobiology indeed in 1924 !</u>



Expected daily change of the reflected light from the earth



Ford, Seager & Turner: Nature 412 (2001) 885

- Assume that the earth's reflected light is completely separated from the Sun's flux !
 - TPF (Terrestrial Planet Finder) in 10 years from now ?
- Periodic change of 10% level due to different reflectivity of land, ocean, forest, and so on
- Cloud is the most uncertain factor: <u>weather forecast</u>

Colors of a Second Earth: estimating the fractional areas of ocean, land and vegetation of Earth-like exoplanets ApJ. 715(2010)866, arXiv:0911.5621 **Colors of a Second Earth. II: Effects of Clouds on Photometric Characterization** of Earth-like Exoplanets ApJ. 738(2011)184, arXiv:1102.3625

Yuka Fujii, H.Kawahara, A.Taruya, Y.Suto (Dept. of Phys., Univ. of Tokyo), S.Fukuda, T.Nakajima (Univ. of Tokyo, Center of climate system research), Edwin Turner (Princeton Univ.) p://www.space.com/scienceastronomy/color-changing-planets-alien-life-100513.html

Colors of our earth





A pale blue dot ? Not really



Simulated photometric light-curves of Earth



Fujii et al. (2010)

Adopted Earth data in March

- Spin inclination = 0 (vernal equinox)
- cloudless

Forward procedure: reflected light model from the earth

- Simulated light-curves of the earth in 7 photometric bands
 - land: BRDF (Bidirectional Reflectance Distribution Function) model from earth-observing satellite Terra/MODIS on 2.5° x2.5° pixels
 - ocean: BRDF model of Nakajima & Tanaka (1983)
 - snow: real data of the month
 - cloud: real data of the day
 - Atmosphere and cloud: radiation transfer solved with rstar6b
- Comparison with the real data observed by EPOXI

Earth observing satellite Trace (Transition Region and Coronal Explorer) + detector Modis (Moderate Resolution Imaging Spectroradiometer)





Simulated light-curves vs. EPOXI data







Inverse procedure: estimation of fractional areas of surface components

- Fitting the EPOXI data to a simplified model (isotropic scattering with ocean, soil, vegetation, snow and cloud)
 - Neglect light from the central star
 - Neglect the spin and orbital rotation during each exposure
 - A simple cloud model with the same optical depth
 τ (=10 fiducially)
 - US standard atmosphere: compositions, pressure and temperature profiles

Cloudless case



Input data

- 5 light-curves using anisotropic scattering (BRDF) model
- 2 week observation of a cloudless Earth at 10 pc away

Inversion assumptions

- Ocean, soil, vegetation and snow only (with atmosphere)
- Isotropic scattering assumed

Results

- Estimated areas (symbols) vs Surface classification data (dashed line)
- Reasonably well reproduced.
- Can identify vegetation !

PCA (principal component analysis)



- 1st eigen vector
 - \Rightarrow soil + vegetation ocean
- 2nd eigen vector
 - i ⇒ vegetation soil ocean – snow



Reconstruction of planetary surface areas with clouds

reflectec

light

emission light Vazquez et al. (2010) Fujii et al. (2010)



Albedo of surface components: isotropic approximation w/o atmosphere



Albedo of surface components: isotropic approximation with atmosphere



Albedo spectra of clouds: model dependence



Fractional areas estimated from EPOXI data



Surface latitude map estimated from EPOXI data





Le Petit Prince: (par Antoine de Saint Exupéry)



Si quelqu'un aime une fleure qui n'existe qu'à un exemplaire dans les millions et les millions d'étoiles, ça suffit pour qu'il soit heureux quand il les regarde. Il se dit: "Ma fleur est là quelque part . . . "